

mind the gap:

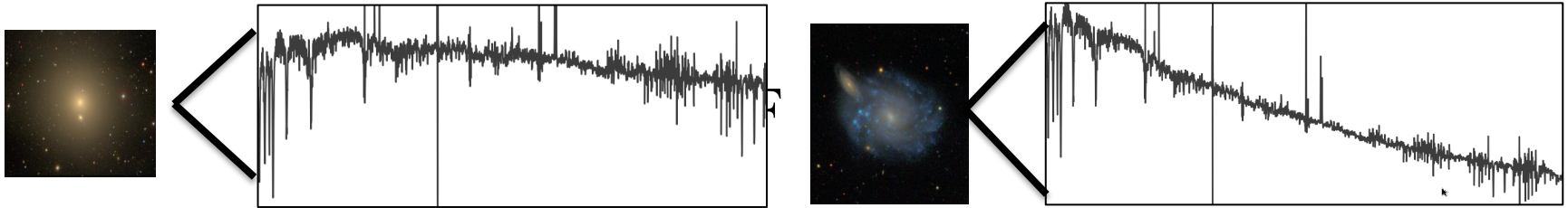
Synergies between resolved star and
integrated light studies of galaxies

Nell Byler
University of Washington

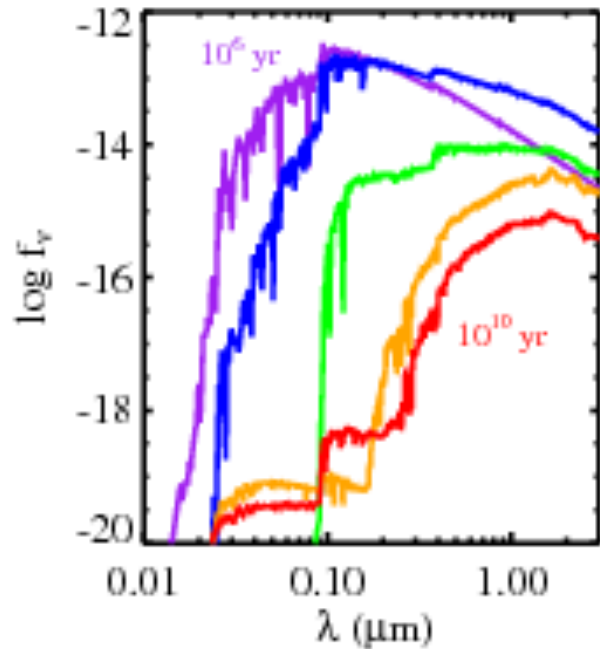
In collaboration with:

Julianne Dalcanton, Ben Williams, Dan Weisz, Ben Johnson, Charlie Conroy

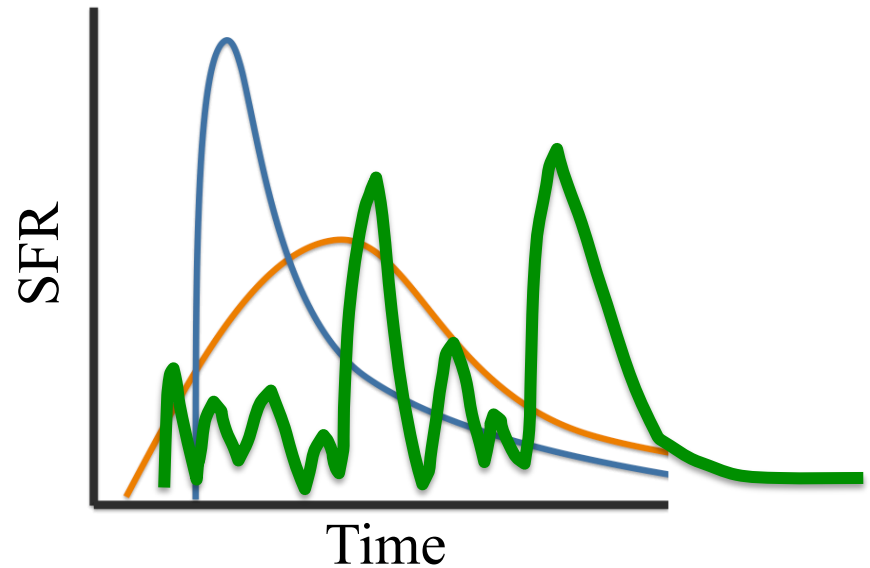
Population Synthesis: Light \rightarrow Astrophysics



SSP

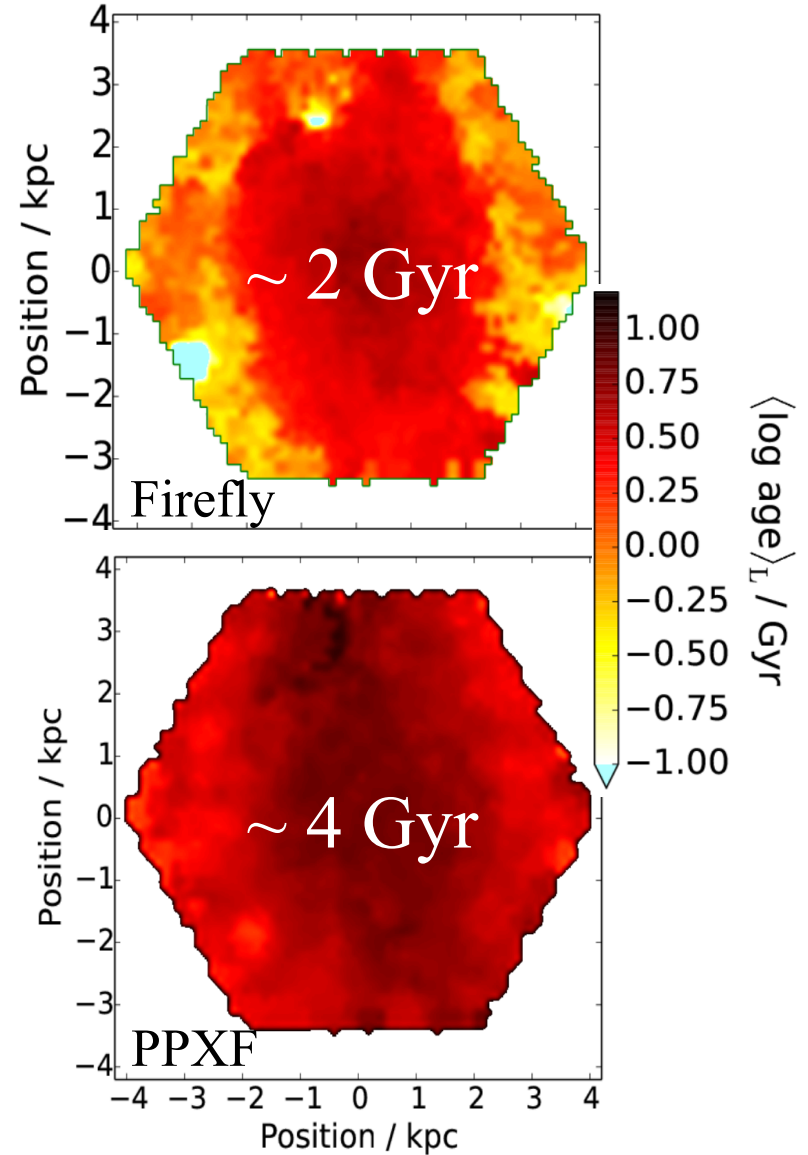
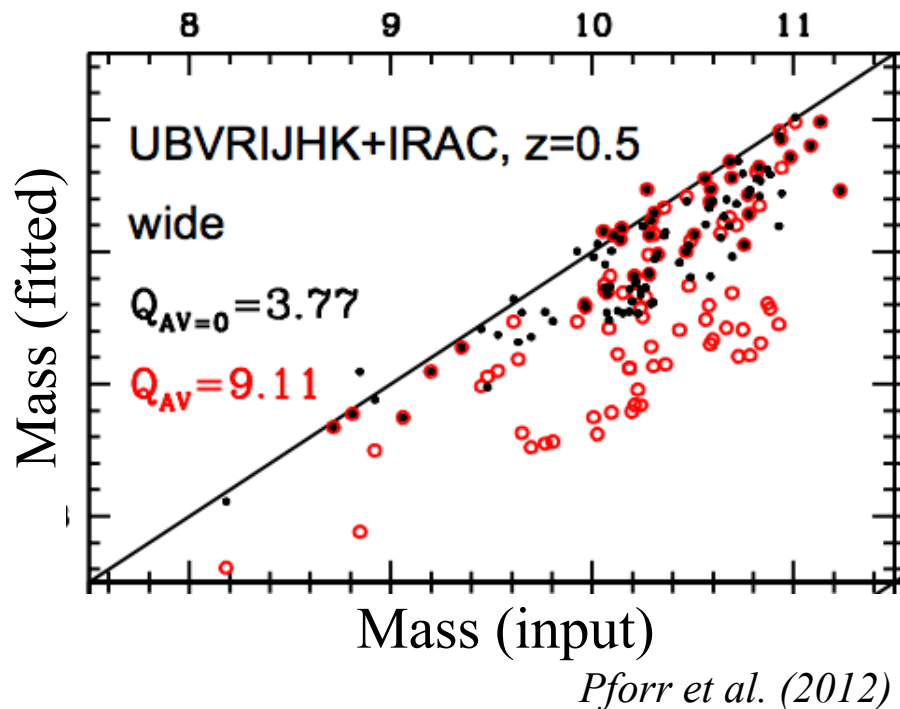


SFH



SFH choice biases derived parameters

M/L , $\langle \text{age} \rangle$, stellar mass, SFR



Wilkinson et al. (2015)

IFU Surveys: New Regime

Variations within individual galaxies of order systematic uncertainties

2" fiber at $z \sim 0.5$

~ 1 kpc at $z = 0.04$



IFU Surveys: New Regime

Variations within individual galaxies of order systematic uncertainties

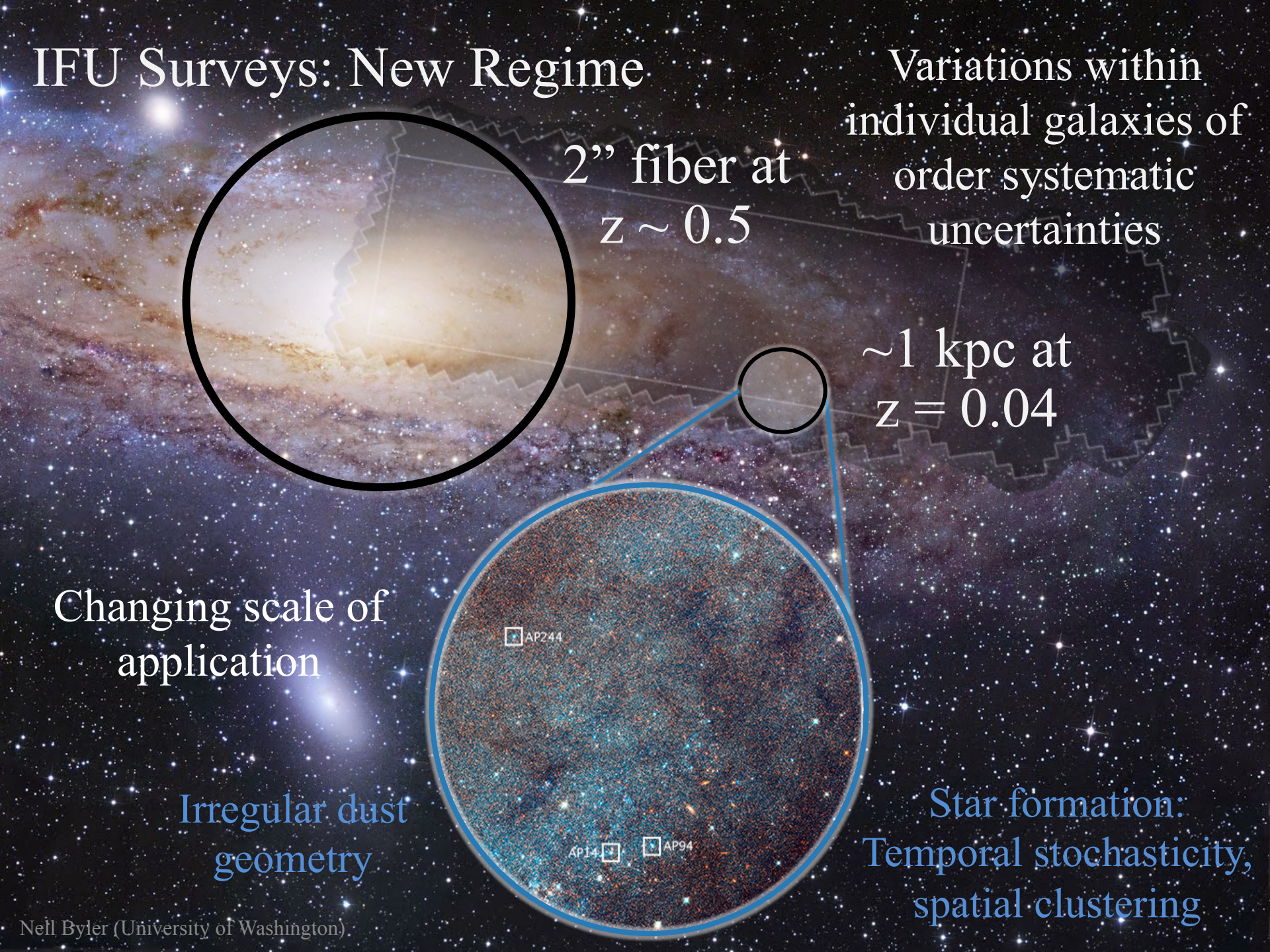
2" fiber at $z \sim 0.5$

~ 1 kpc at $z = 0.04$

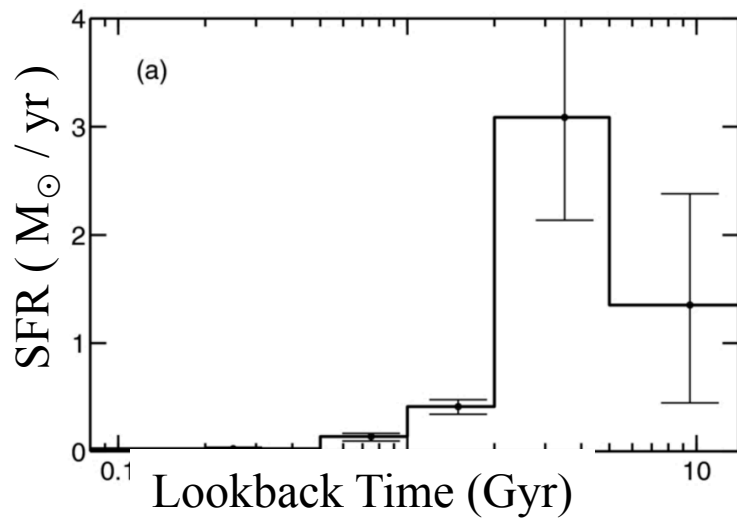
Changing scale of application

Irregular dust geometry

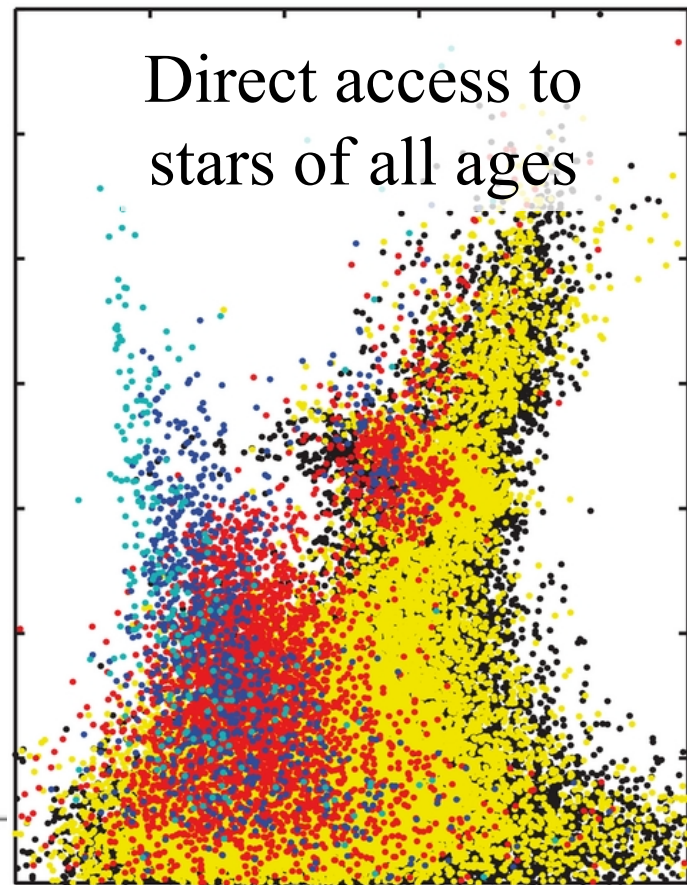
Star formation: Temporal stochasticity, spatial clustering



Resolved Stars: “gold standard” SFHs



Mag (F814W)



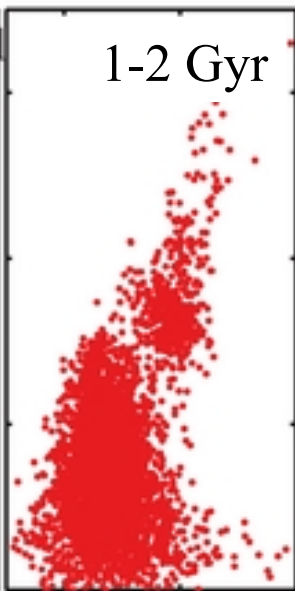
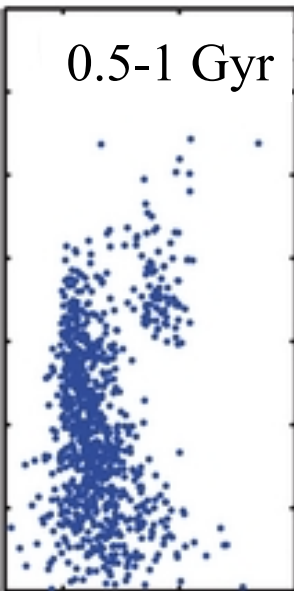
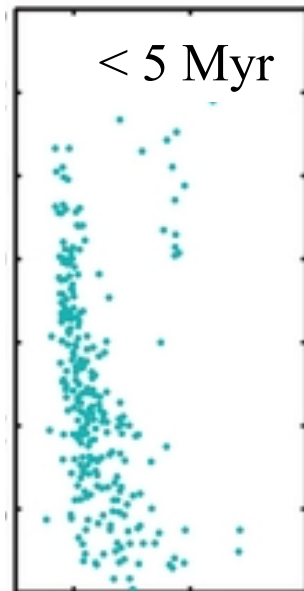
< 5 Myr

0.5-1 Gyr

1-2 Gyr

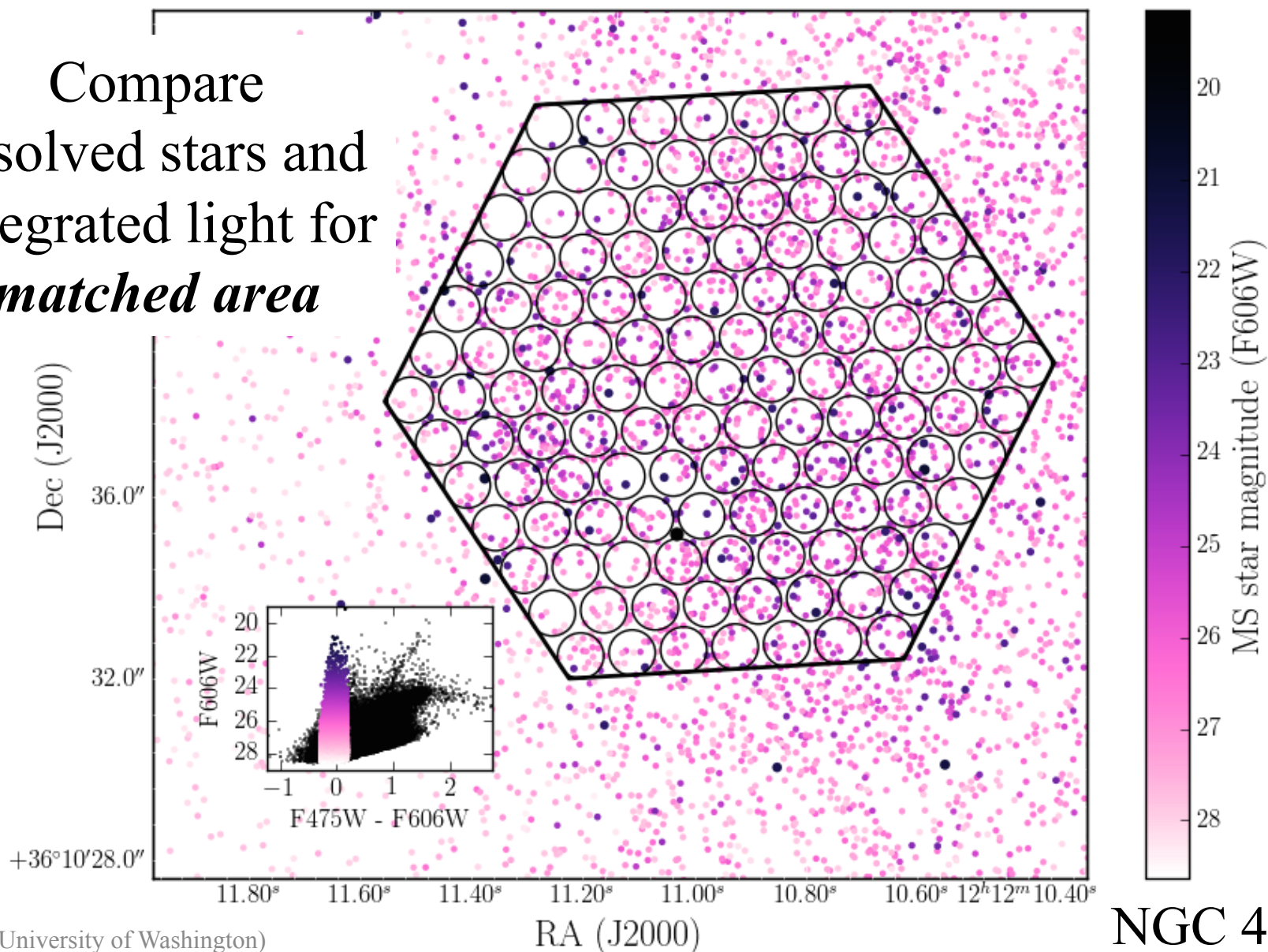
2-5 Gyr

Color (F606W - F814W)



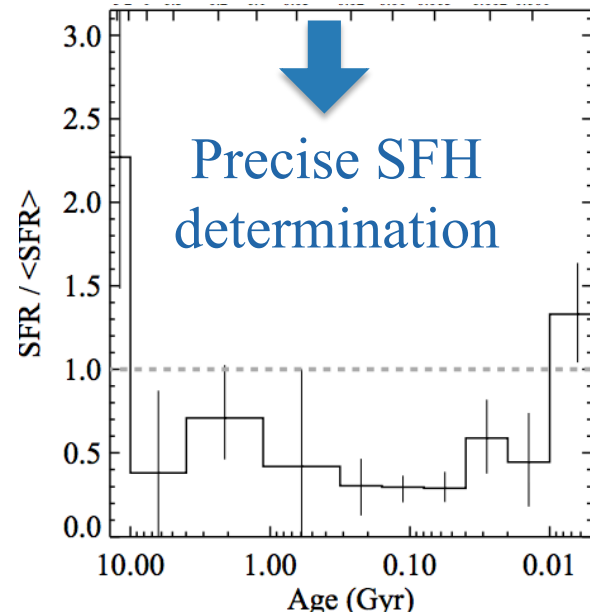
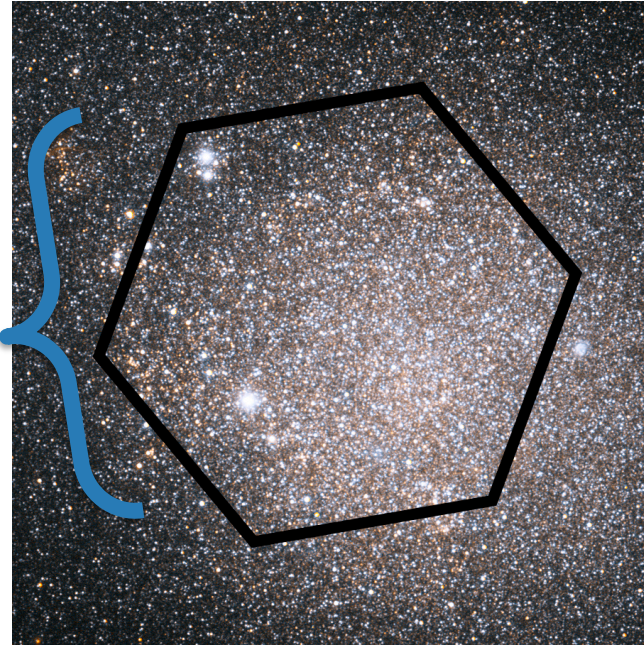
Empirical Calibration: HST + MaNGA

Compare
resolved stars and
integrated light for
matched area

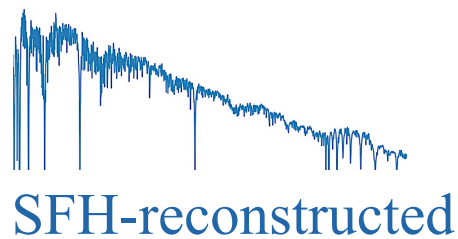


Aperture-Matched Comparison of Light

Resolved Stars

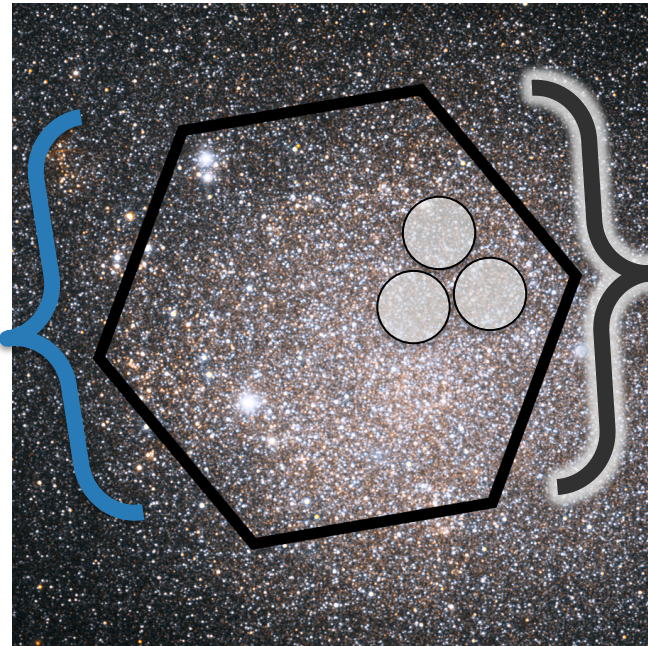


Precise SFH
determination

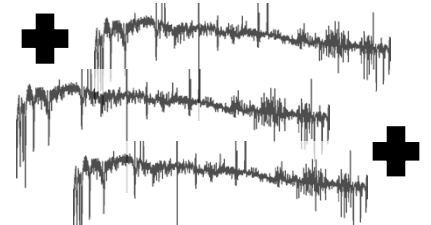


Aperture-Matched Comparison of Light

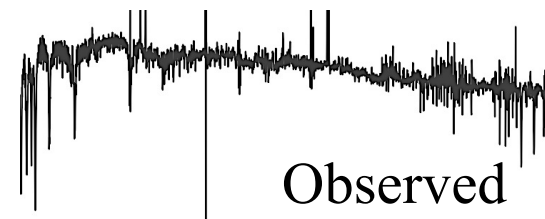
Resolved Stars



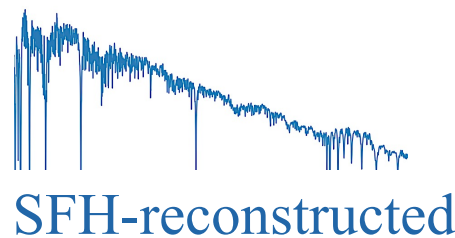
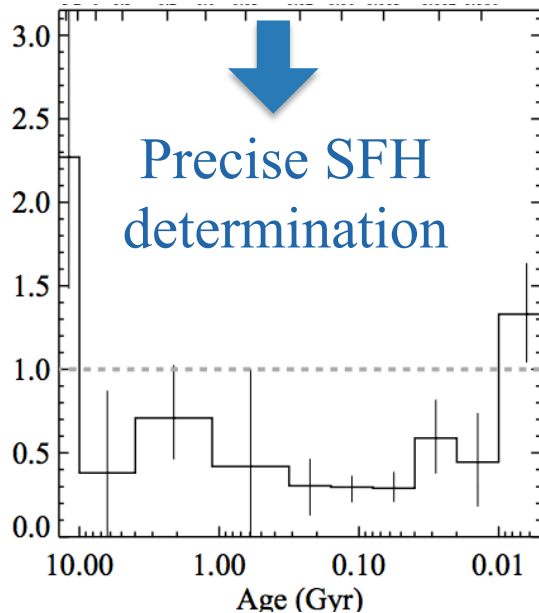
Integrated Light



Sum MaNGA spectra
over matched area



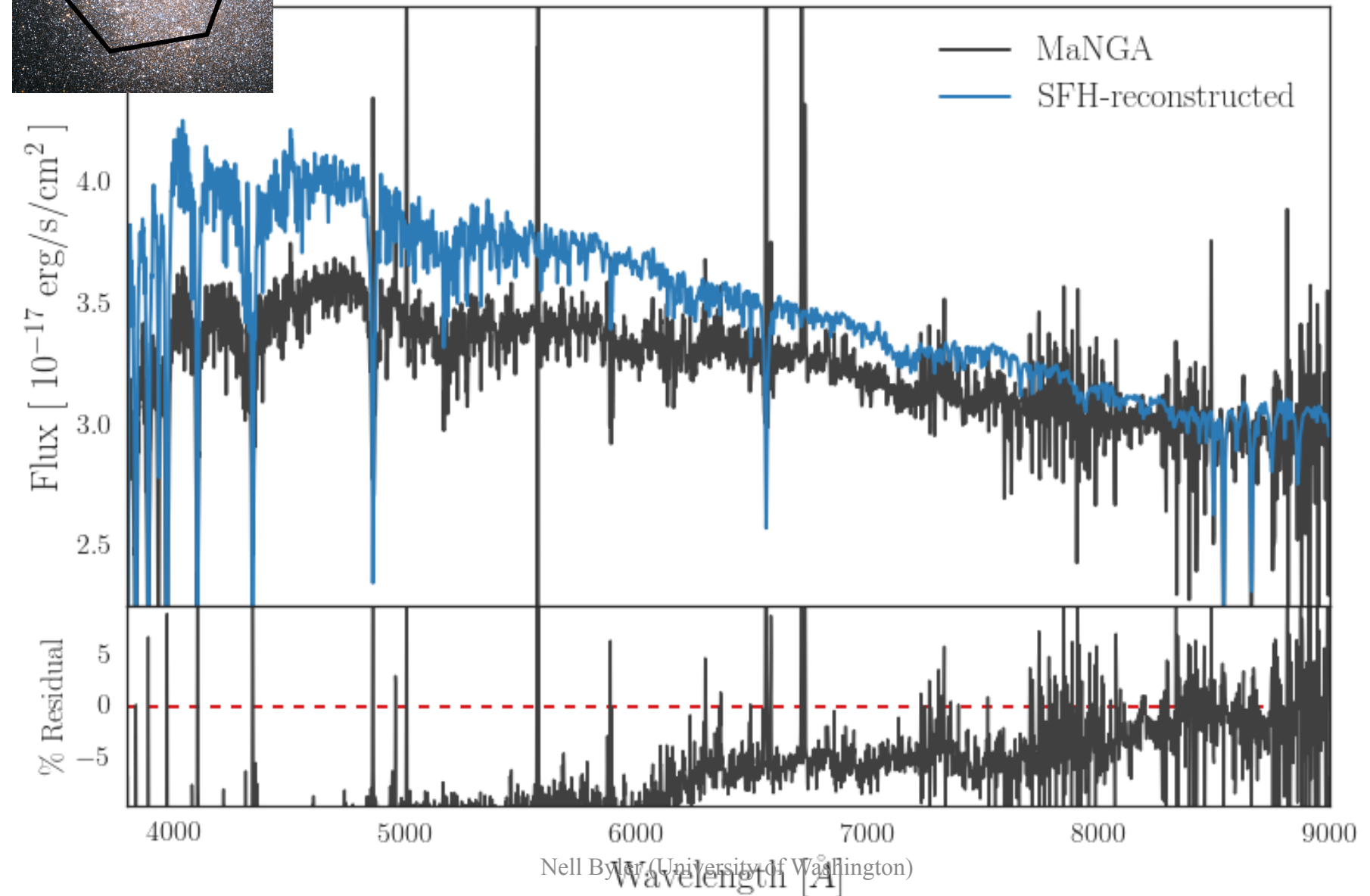
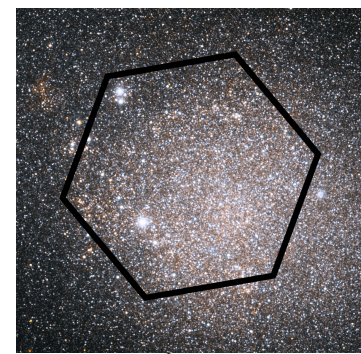
Observed



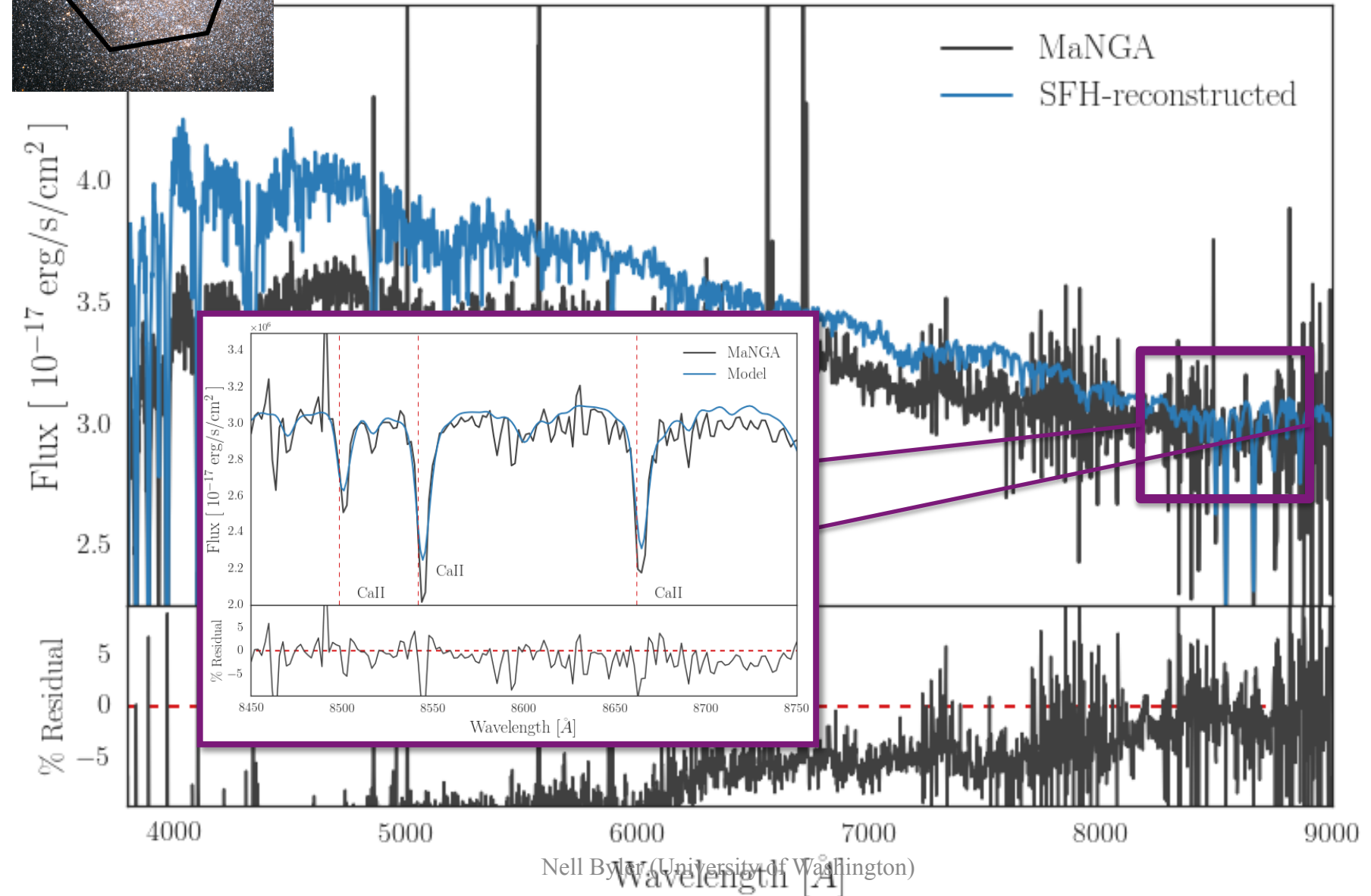
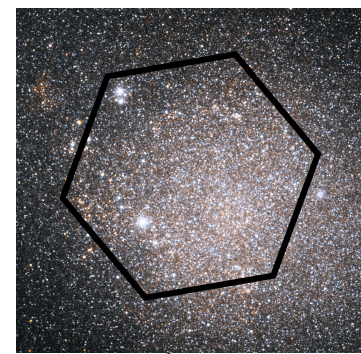
SFH-reconstructed

*Directly measure contribution
of stars to total light*

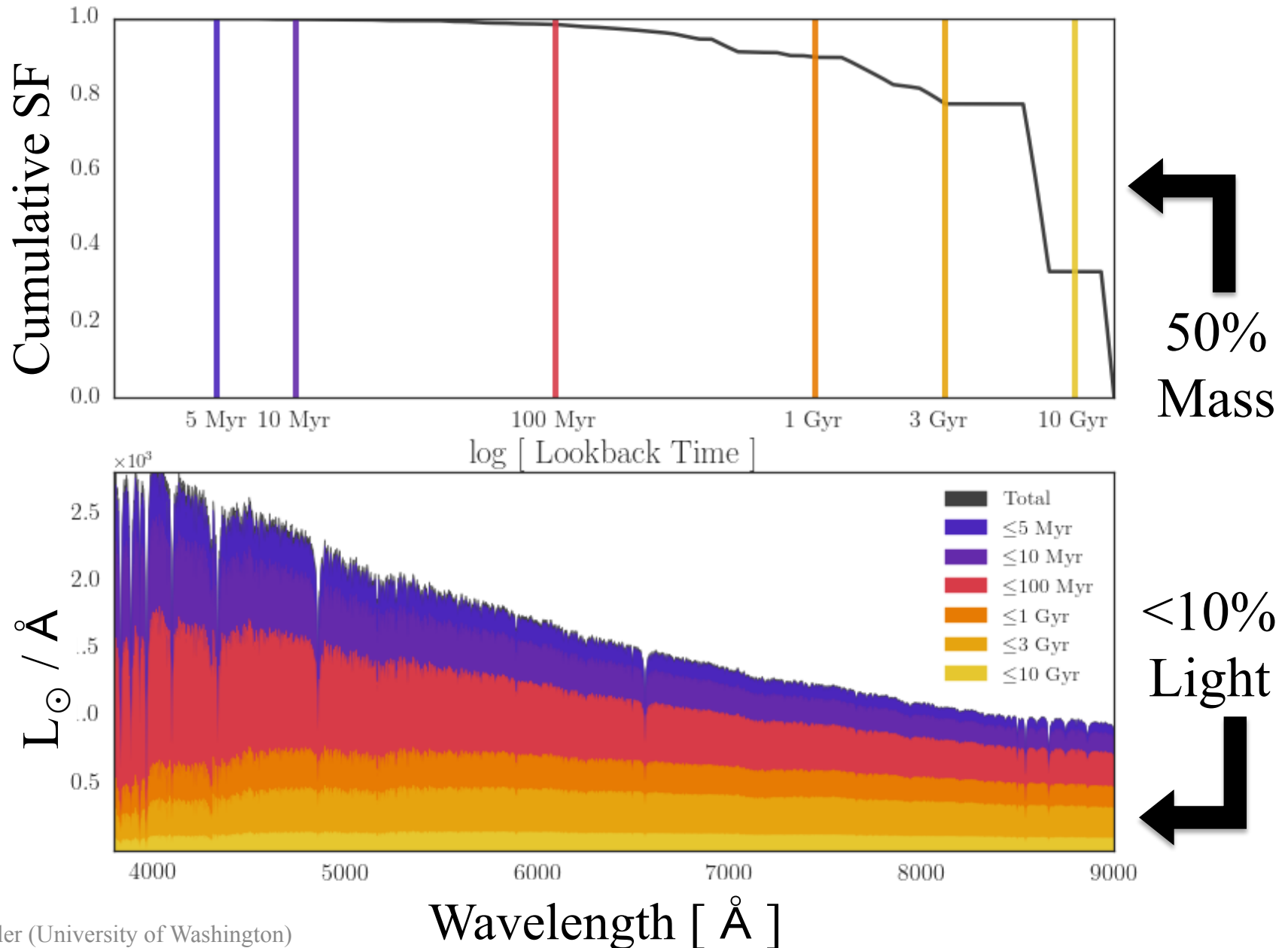
Aperture-Matched Comparison



Aperture-Matched Comparison



Assess sensitivity of mass-weighted properties



Empirical calibration: Test your code here!

