

SPECTRAL EVOLUTION OF GALAXIES. II AN ATLAS OF *IUE* SPECTRA OF LATE TYPE STARS AND NEARBY EARLY TYPE STELLAR SYSTEMS

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RESUMEN

Se presentan en forma gráfica y tabular espectros de baja dispersión de estrellas tardías, cúmulos globulares y galaxias elípticas obtenidos con el satélite *IUE*. Para aquellos tipos de objetos con más de un espectro observado, se calculan espectros promedios.

ABSTRACT

Low dispersion *IUE* spectra are presented in both pictorial and tabular form for late type stars, globular clusters, and elliptical galaxies. Average spectra are computed for types of objects with more than one observed spectrum.

Key words: LATE TYPE STARS—GLOBULAR CLUSTERS—ELLIPTICAL GALAXIES—ULTRA-VIOLET SPECTRA

I. INTRODUCTION

The *International Ultraviolet Explorer* satellite (*IUE*) was used to obtain ultraviolet spectra of late type stars and early type galaxies. These spectra were needed in order to complete an extensive project that involves the study of the spectral evolution of galaxies in the wavelength range from the ultraviolet to the infrared (Bruzual 1981, 1983a, 1983b, 1983c). The stellar data were required to complete the library of stellar spectra used in the synthesis. For spectral types earlier than G2 the *OAO-2* data of Code and Meade (1979) were used. The galaxy data were needed to define the typical early type galaxy UV spectrum to be compared with the results of the evolutionary synthesis. Due to the apparent faintness of the sources, the low dispersion mode of the *IUE* was used to maximize the signal-to-noise ratio and to make the exposure times reasonable. This mode provides enough resolution ($\sim 5 \text{ \AA}$) for the synthesis program.

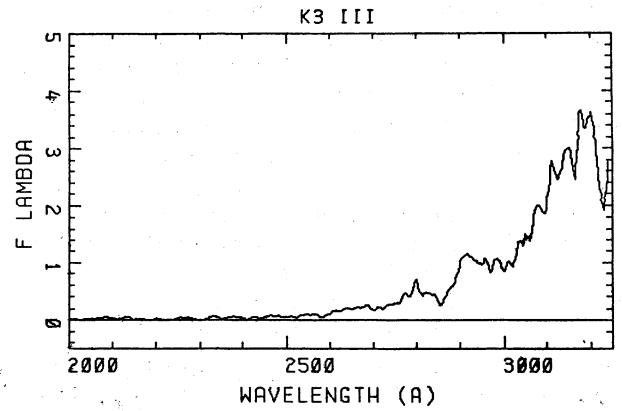
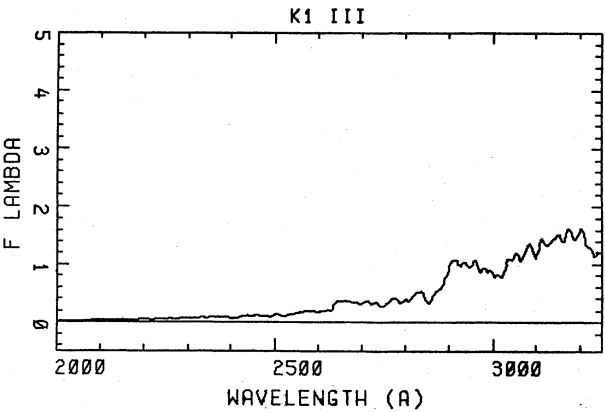
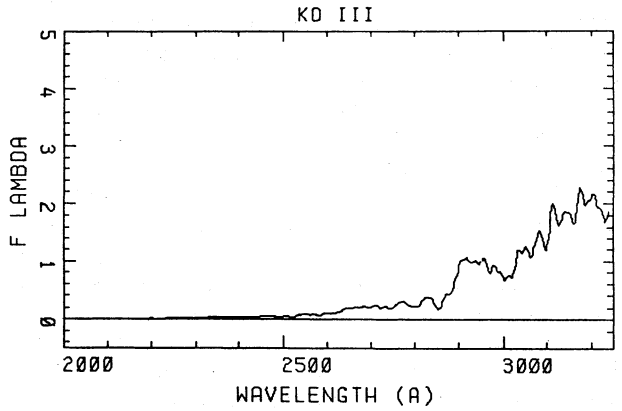
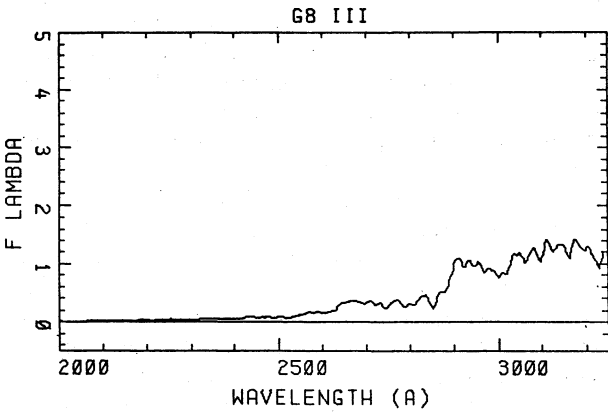
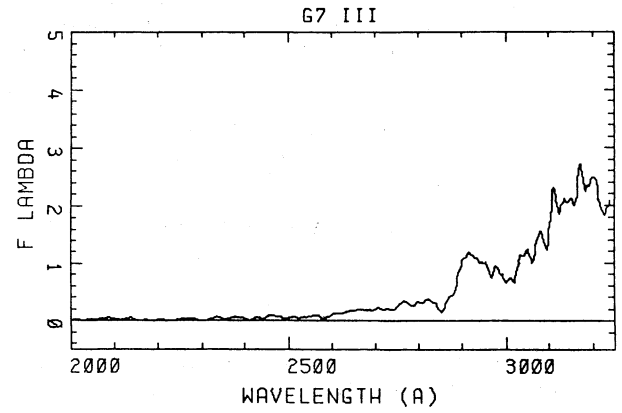
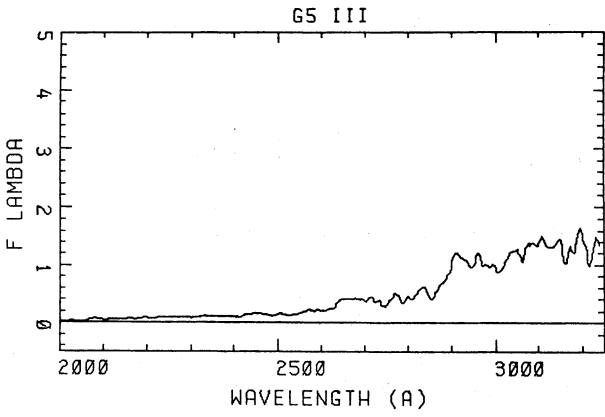
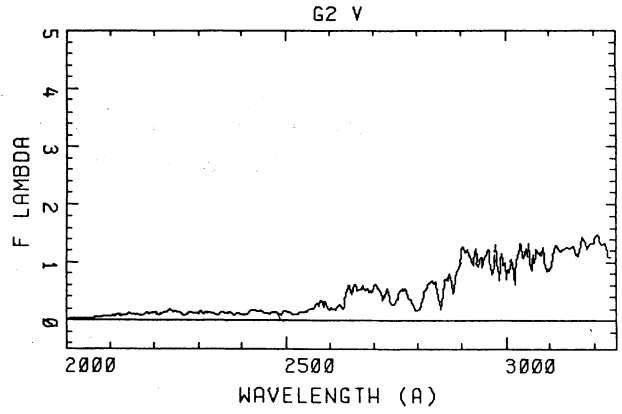
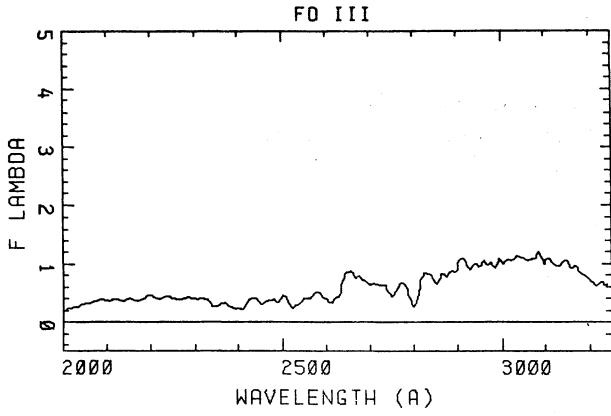
Besides the data acquired by the author, several *IUE* spectra of both stars and stellar systems were provided by the National Space Science Data Center (see Tables 1 and 2). While collecting these data it became apparent that there was a need for a uniform presentation of these spectra, similar to that given by Code and Meade (1979) for the *OAO-2* stellar data. Much of the data are dispersed throughout the literature, but the spectra have not been presented in a consistent and coherent form. In this paper I present such a compilation. All the spectra are displayed in both pictorial and tabular form in a common wavelength scale. Section II describes the details of the *IUE* data reduction performed by the author. The results are presented in section III.

II. DATA REDUCTION

The final reduction of the spectra was performed based on the files provided by the *IUE* data reduction facilities. Each spatially resolved spectral segment (line-by-line spectrum file) provided by NASA or ESA was plotted individually, as a spectrum F_λ versus λ . By visual inspection those segments with the highest signal-to-noise, i.e., with the spectral shape clearly visible, were chosen as defining the source spectrum. The spectra of extended sources typically included from eight to ten segments; in some of the exposures, the source was off-center in the large aperture. Special care was taken used to define the background level. Segments were chosen to cover most of the camera, approaching as close as possible to the source spectrum, without including segments with recognizable spectral features. The segments defining the background were averaged and smoothed, then normalized to the same area as the spectrum, and finally subtracted. At this stage the net counts were calibrated as usual (Bohlin and Holm 1980). This procedure improved the final spectrum considerably, especially below 2500 \AA , over the standard reduction. This can be seen by comparing Figure 2a with Figure 1 of Bruzual and Spinrad (1981). The latter was constructed with the spectra reduced by the NASA facilities. The remaining high frequency noise in the final spectra should not be taken as spectral features. Quite often these apparent emission peaks represent particles hitting the detector at any time during the exposure.

III. RESULTS

Table 1 lists the stars, spectral type, and LWR image



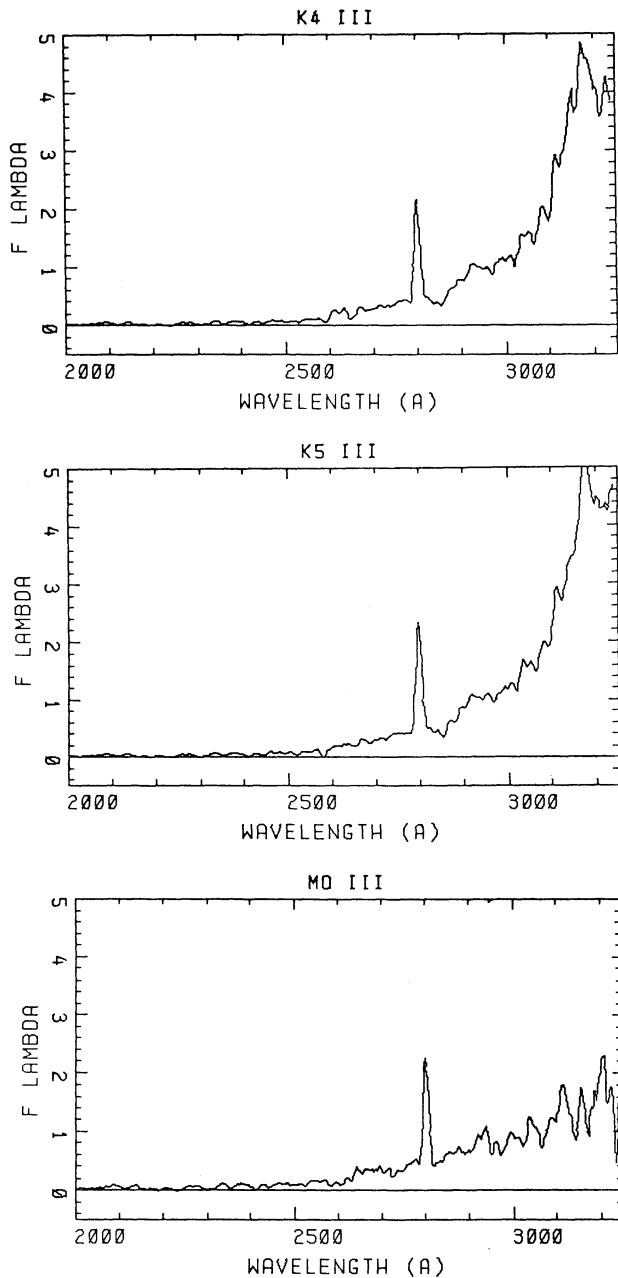


Fig. 1. (a,b,c,d,e,f,g,h,i,j,k) Ultraviolet spectra obtained with the *IUE* satellite for stars of spectral types as indicated in Table 1 and on top of each figure. For the spectral types with more than one entry in Table 1, an average spectrum was determined. All the spectra have been normalized to $F_{\lambda} = 1$ at $\lambda = 2940$ Å.

number for which *IUE* data were obtained. For spectral types of stars with more than one entry in this table an average spectrum was used after normalization at $\lambda = 2940$ Å. The spectra in the range from 2000 to 3240 Å are displayed in Figure 1 as plots of F_{λ} versus λ .

Table 2 indicates the identification, color excess, *IUE* image number, and exposure time for the elliptical gal-

TABLE 1

STARS OBSERVED WITH <i>IUE</i>		
Sp. Type	Star	LWR Image No.
F0 III	γ Tuc	4688
G2 V	{ Sun	(*)
...	{ 16 Cyg A	1888, 1889
G5 III	ι Tuc	4689
G7 III	λ^2 Tuc	4690
G8 III	{ ϵ And	4670
...	{ η And	4671
...	{ λ And	4666
K0 III	{ α Cas	4646
...	{ 3 And	4667, 4668
K1 III	β Cet	4672, 4687
K3 III	δ And	4669
K4 III	ϕ^3 Cet	4686
K5 III	δ Psc	4684, 4685
M0 III	β And	4683

* The data for the sun was taken from Broadfoot (1972) and Arvensen *et al.* (1969).

axies and globular clusters for which *IUE* spectra were available. The reddening law (analytical approximation from Seaton (1979) was used to correct the observed *UV* spectra. As in the optical region, the degree of similarity between the *UV* spectra of these systems in the 2000 to 3200 Å region is very high, and it is possible to define an average *UV* spectrum for each type of stellar system considered. Different spectra were assigned weights in proportion to the observing time and in accord with spectral quality, as given in Table 3.

Table 4 lists *IUE* stellar spectra in numerical form. All the spectra presented in this atlas have been normalized to $F_{\lambda} = 1$ at $\lambda = 2940$ Å. The resulting average spectra of stellar systems in the range from 2000 to 3200 Å are shown in Figure 2 and listed in Table 5.

Finally, Figure 3 shows the average spectrum of M 31 and M 32, as well as the spectra of the globular clusters NGC 6752 and 47 Tuc in the range from 1200 to 3200 Å. These were the only systems observed down to 1200 Å. The fluxes are listed in Table 6. The fluxes from 1940 to 1995 Å are listed as zero; in this region there is no good overlap between the *IUE* SWP and LWR cameras.

IV. CONCLUDING REMARKS

The *IUE* spectra of late type stars, globular clusters, and elliptical galaxies presented in section III are expected to be representative of objects of their kinds. These spectra can be used to estimate *UV* luminosities or for spectral feature identification purposes.

None of the galaxies included in this work show the rapid increase shortward of 2000 Å that has been observed in NGC 4486 (Bertola *et al.* 1980; Perola and Tarenghi 1980) and NGC 4649 (Bertola, Capaccioli, and Oke 1982). These two galaxies are thus "peculiar" with

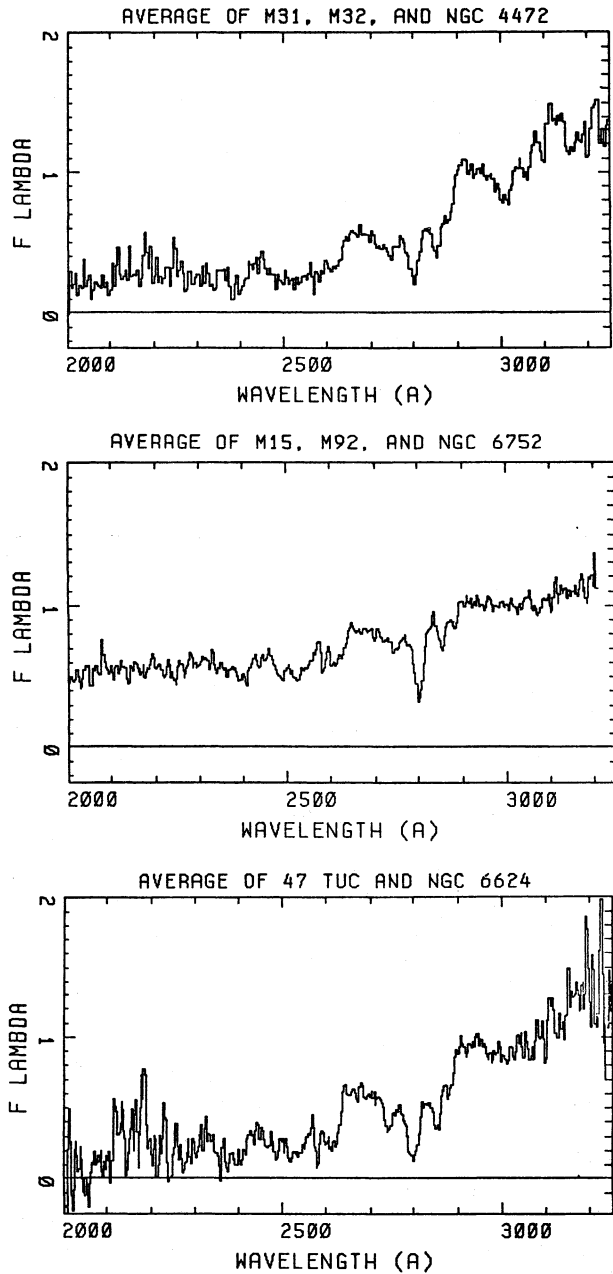


Fig. 2. Average spectral energy distributions in the range from 2000 to 3200 Å for: (a) M 31, M 32, and NGC 4472, (b) M 15, M 92, and NGC 6752, and (c) 47 Tuc and NGC 6624. Each spectrum was normalized at 2940 Å and weighted according to exposure time and spectral quality as indicated in Table 3. The high frequency structure shortward of 2600 Å is not real, being most likely due to detector noise, or particles reaching the camera during the exposures. The narrow "absorption feature" around 2580 Å seen in some of the spectra is produced by a reseau mark in the camera.

respect to M31 and M32. Nor do any of these galaxies show emission lines as seen in the spectrum of M81 displayed by Peimbert and Torres-Peimbert (1981) and Bruzual, Peimbert, and Torres-Peimbert (1982).

TABLE 2

STELLAR SYSTEMS OBSERVED WITH IUE

Object	Type	E(B-V)	Image No.	Exposure (hrs.)
M 31 (bulge)	E - gal	0.10	SWP 3520	8.3
			LWR 3088	5.0
			LWR 4665 (*)	1.0
			LWR 4682 (*)	2.5
M 32	E - gal	0.10	SWP 3545	5.0
			LWR 3111	6.3
			LWR 4644 (*)	3.0
			LWR 4645 (*)	2.5
NGC 4472	E - gal	0.02	LWR 4520	3.0
			LWR 4521	3.5
			LWR 4664 (*)	6.0
			LWR 4681 (*)	4.5
M 15	G. C.	0.09	LWR 1457	0.2
			LWR 1820	1.0
			LWR 1821	1.2
M 92	G. C.	0.02	LWR 1448	0.5
			LWR 1881	0.7
			LWR 1882	0.7
NGC 6752	G. C.	0.04	SWP 1531	0.7
			SWP 1532	0.7
			LWR 1482	0.5
			LWR 1839	0.5
47 Tuc	G. C.	0.06	SWP 1510	3.0
			LWR 1461	0.5
NGC 6624	G. C.	0.30	LWR 1458	1.5

* Data obtained by H. Spinrad and the author.

TABLE 3

AVERAGE S. E. D. OF DIFFERENT STELLAR SYSTEMS

Object	Range (Å)	Ingredients	Weight
E - gal	1250-3200	M 31 (bulge)	0.50
		M 32	0.50
	2000-3200	M 31 (bulge)	0.30
		M 32	0.40
G. C. (metal poor)	2000-3200	NGC 4472	0.30
		M 15	0.55
		M 92	0.30
G. C. (metal rich)	2000-3200	NGC 6752	0.15
		47 Tuc	0.75
		NGC 6624	0.25

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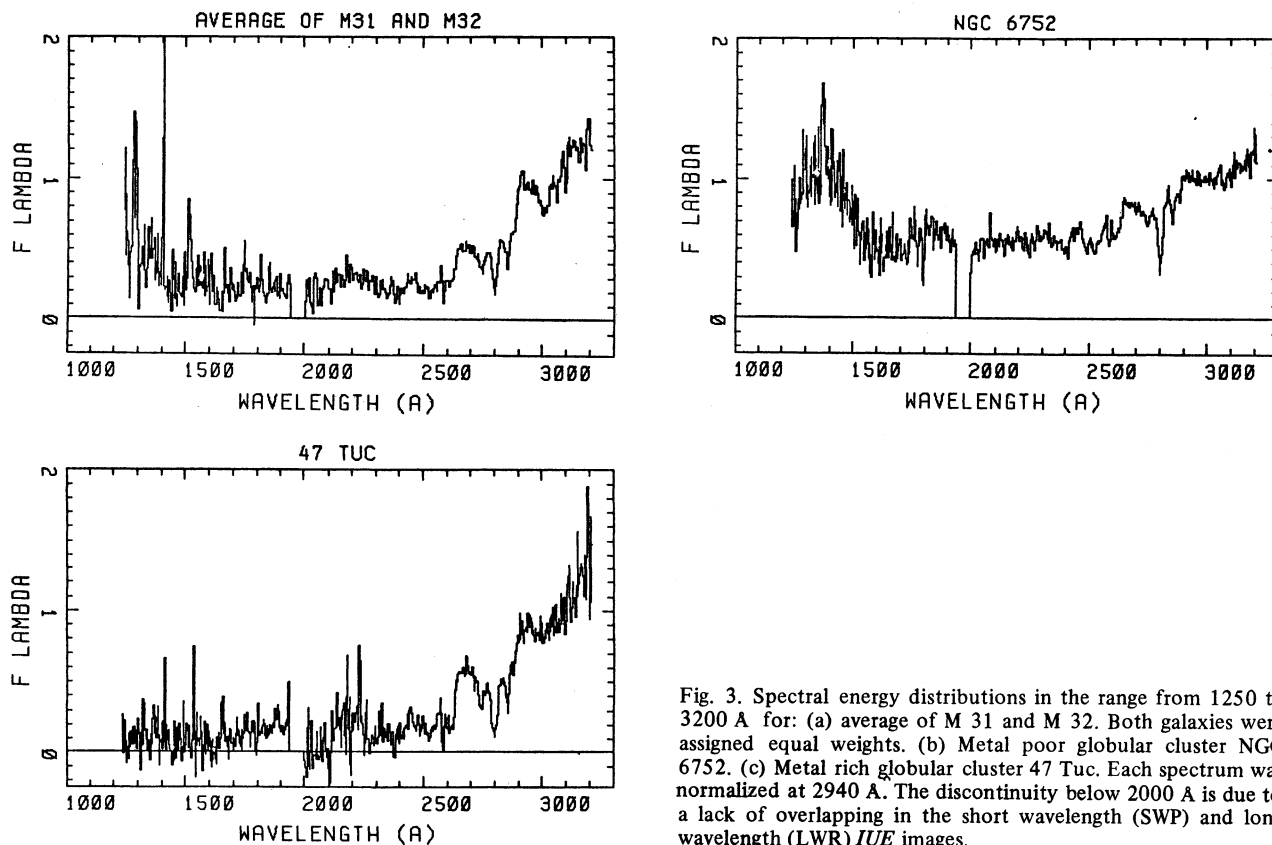


Fig. 3. Spectral energy distributions in the range from 1250 to 3200 Å for: (a) average of M 31 and M 32. Both galaxies were assigned equal weights. (b) Metal poor globular cluster NGC 6752. (c) Metal rich globular cluster 47 Tuc. Each spectrum was normalized at 2940 Å. The discontinuity below 2000 Å is due to a lack of overlapping in the short wavelength (SWP) and long wavelength (LWR) IUE images.

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TABLE 5

IUE SPECTRA OF STELLAR SYSTEMS (2000-3200 Å)

$\lambda(\text{Å})$	E-GAL F_{λ}	MP-GC F_{λ}	MR-GC F_{λ}	$\lambda(\text{Å})$	E-GAL F_{λ}	MP-GC F_{λ}	MR-GC F_{λ}	$\lambda(\text{Å})$	E-GAL F_{λ}	MP-GC F_{λ}	MR-GC F_{λ}	$\lambda(\text{Å})$	E-GAL F_{λ}	MP-GC F_{λ}	MR-GC F_{λ}	$\lambda(\text{Å})$	E-GAL F_{λ}	MP-GC F_{λ}	MR-GC F_{λ}				
2000	-0.109	0.203	-0.098	2240	0.264	0.622	0.023	2480	0.227	0.671	0.270	2720	0.492	0.847	0.597	2960	1.076	1.043	0.985				
2005	0.140	0.410	-0.207	2245	0.460	0.569	0.014	2485	0.252	0.660	0.150	2725	0.472	0.837	0.571	2965	1.001	1.034	0.911				
2010	0.234	0.484	-0.290	2250	0.491	0.541	0.023	2490	0.247	0.644	0.151	2730	0.504	0.839	0.525	2970	0.983	0.992	0.868				
2015	0.202	0.436	0.508	2255	0.319	0.597	0.229	2495	0.224	0.634	0.222	2735	0.472	0.833	0.428	2975	1.006	0.998	0.921				
2020	0.278	0.556	-0.128	2260	0.310	0.635	0.403	2500	0.227	0.682	0.267	2740	0.442	0.802	0.341	2980	1.000	1.009	0.861				
2025	0.143	0.505	-0.079	2265	0.365	0.653	0.137	2505	0.306	0.717	0.187	2745	0.410	0.790	0.363	2985	0.927	1.018	0.910				
2030	0.204	0.529	0.289	2270	0.283	0.633	0.250	2510	0.241	0.674	0.291	2750	0.439	0.808	0.386	2990	0.923	1.017	0.920				
2035	0.183	0.535	0.047	2275	0.161	0.692	0.041	2515	0.224	0.664	0.155	2755	0.483	0.822	0.476	2995	0.909	1.006	0.963				
2040	0.342	0.577	0.232	2280	0.228	0.709	0.117	2520	0.246	0.615	0.123	2760	0.490	0.834	0.487	3000	0.834	1.018	0.853				
2045	0.216	0.541	-0.002	2285	0.285	0.591	0.189	2525	0.185	0.602	0.151	2765	0.492	0.848	0.487	3005	0.808	1.015	0.895				
2050	0.256	0.523	-0.126	2290	0.227	0.614	0.161	2530	0.241	0.655	0.186	2770	0.560	0.862	0.527	3010	0.853	1.018	0.866				
2055	0.146	0.528	-0.120	2295	0.251	0.669	0.204	2535	0.214	0.662	0.191	2775	0.533	0.847	0.462	3015	0.814	0.994	0.857				
2060	0.147	0.557	-0.208	2300	0.279	0.637	0.268	2540	0.231	0.675	0.183	2780	0.497	0.883	0.403	3020	0.851	1.001	0.956				
2065	0.222	0.534	0.044	2305	0.264	0.659	0.153	2545	0.264	0.699	0.189	2785	0.432	0.830	0.316	3025	0.919	0.998	0.933				
2070	0.210	0.546	0.102	2310	0.215	0.635	0.236	2550	0.260	0.722	0.236	2790	0.369	0.731	0.190	3030	1.003	1.017	0.864				
2075	0.191	0.560	0.096	2315	0.295	0.663	0.284	2555	0.271	0.737	0.258	2795	0.293	0.616	0.171	3035	1.063	1.028	1.041				
2080	0.207	0.605	0.116	2320	0.288	0.661	0.208	2560	0.336	0.752	0.320	2800	0.216	0.535	0.131	3040	0.062	1.041	0.986				
2085	0.233	0.628	0.128	2325	0.321	0.668	0.365	2565	0.280	0.786	0.323	2805	0.274	0.598	0.206	3045	1.119	1.026	1.062				
2090	0.205	0.601	0.048	2330	0.209	0.666	0.268	2570	0.156	0.798	0.467	2810	0.393	0.720	0.244	3050	1.039	1.014	0.883				
2095	0.151	0.546	0.124	2335	0.208	0.627	0.322	2575	0.304	0.823	0.231	2815	0.475	0.848	0.450	3055	1.004	1.038	1.073				
2100	0.156	0.515	0.162	2340	0.163	0.634	0.272	2580	0.276	0.680	0.079	2820	0.588	0.918	0.558	3060	0.994	0.995	0.942				
2105	0.356	0.472	0.191	2345	0.202	0.621	0.321	2585	0.254	0.508	0.115	2825	0.610	0.904	0.546	3065	1.035	0.973	0.874				
2110	0.216	0.490	-0.032	2350	0.314	0.638	0.183	2590	0.304	0.644	0.331	2830	0.609	0.912	0.538	3070	1.149	0.985	0.961				
2115	0.362	0.589	0.498	2355	0.312	0.592	0.167	2595	0.345	0.736	0.332	2835	0.615	0.907	0.533	3075	1.223	1.003	0.879				
2120	0.401	0.574	0.523	2360	0.315	0.634	-0.021	2600	0.368	0.723	0.247	2840	0.585	0.897	0.520	3080	1.319	1.063	1.164				
2125	0.264	0.538	0.336	2365	0.269	0.615	0.230	2605	0.328	0.693	0.274	2845	0.493	0.843	0.459	3085	1.252	1.067	1.031				
2130	0.255	0.578	0.338	2370	0.321	0.644	0.127	2610	0.309	0.728	0.222	2850	0.451	0.817	0.363	3090	1.229	1.048	1.034				
2135	0.304	0.580	0.560	2375	0.218	0.609	0.080	2615	0.336	0.752	0.208	2855	0.435	0.821	0.394	3095	1.156	1.058	1.020				
2140	0.254	0.517	0.349	2380	0.104	0.581	0.117	2620	0.330	0.775	0.274	2860	0.539	0.849	0.474	3100	1.106	1.028	0.878				
2145	0.478	0.529	0.025	2385	0.128	0.596	0.176	2625	0.355	0.772	0.251	2865	0.642	0.903	0.647	3105	1.297	1.041	1.166				
2150	0.267	0.518	0.182	2390	0.225	0.611	0.140	2630	0.385	0.786	0.332	2870	0.687	0.917	0.674	3110	1.424	1.077	1.315				
2155	0.275	0.543	0.362	2395	0.213	0.599	0.103	2635	0.448	0.830	0.481	2875	0.669	0.914	0.678	3115	1.528	1.143	1.315				
2160	0.289	0.560	0.357	2400	0.154	0.580	0.147	2640	0.515	0.889	0.633	2880	0.681	0.893	0.616	3120	1.410	1.112	1.213				
2165	0.256	0.558	0.465	2405	0.189	0.600	0.134	2645	0.553	0.896	0.685	2885	0.736	0.893	0.687	3125	1.414	1.107	1.025				
2170	0.233	0.519	0.344	2410	0.236	0.586	0.139	2650	0.531	0.911	0.611	2890	0.820	0.950	0.766	3130	1.392	1.091	1.121				
2175	0.398	0.602	0.318	2415	0.252	0.631	0.210	2655	0.582	0.897	0.614	2895	0.947	0.991	0.888	3135	1.429	1.108	1.093				
2180	0.566	0.582	0.757	2420	0.370	0.674	0.203	2660	0.597	0.887	0.634	2900	1.018	0.999	0.910	3140	1.433	1.113	1.143				
2185	0.434	0.624	0.802	2425	0.343	0.701	0.321	2665	0.583	0.900	0.641	2905	1.061	0.990	0.965	3145	1.419	1.160	1.188				
2190	0.362	0.688	0.734	2430	0.328	0.743	0.346	2670	0.564	0.853	0.585	2910	1.087	0.993	0.985	3150	1.238	1.080	1.302				
2195	0.428	0.730	0.216	2435	0.382	0.723	0.319	2675	0.611	0.865	0.617	2915	1.115	0.990	0.964	3155	1.190	1.112	1.251				
2200	0.241	0.629	0.287	2440	0.341	0.673	0.349	2680	0.589	0.872	0.702	2920	1.114	1.018	0.894	3160	1.163	1.120	1.373				
2205	0.343	0.602	0.088	2445	0.373	0.713	0.396	2685	0.580	0.878	0.632	2925	1.018	0.997	0.983	3165	1.203	1.184	1.329				
2210	0.285	0.613	0.315	2450	0.449	0.729	0.291	2690	0.571	0.892	0.596	2930	1.052	1.003	0.975	3170	1.195	1.164	1.436				
2215	0.265	0.601	0.015	2455	0.399	0.726	0.366	2695	0.563	0.885	0.608	2935	1.068	1.011	0.999	3175	1.281	1.109	1.339				
2220	0.221	0.540	0.306	2460	0.319	0.789	0.278	2700	0.536	0.862	0.601	2940	1.000	1.000	1.000	3180	1.287	1.081	1.239				
2225	0.263	0.587	0.176	2465	0.295	0.756	0.236	2705	0.589	0.868	0.638	2945	1.031	0.992	1.052	3185	1.256	1.129	1.435				
2230	0.333	0.636	0.549	2470	0.314	0.707	0.240	2710	0.541	0.849	0.633	2950	1.057	0.974	1.017	3190	1.300	1.223	1.346				
2235	0.298	0.602	0.432	2475	0.280	0.661	0.305	2715	0.491	0.854	0.578	2955	1.025	1.008	1.017	3195	1.389	1.117	1.824				
																				3200	1.136	1.171	1.291

2450	0.328	0.145	0.110	0.030	0.045	0.033	0.085	0.050	0.017	0.046	0.053	3045	1.105	1.235	1.246	1.195	1.133	1.168	1.077	1.353	1.531	1.377	1.080
2451	0.358	0.154	0.135	0.049	0.073	0.036	0.098	0.057	0.036	0.050	0.066	3050	1.062	1.279	1.289	1.238	1.177	1.254	1.191	1.466	1.599	1.464	1.599
2452	0.403	0.166	0.138	0.051	0.079	0.040	0.106	0.058	0.059	0.050	0.089	3055	1.104	1.107	1.189	1.149	1.129	1.210	1.187	1.455	1.587	1.466	1.627
2453	0.416	0.156	0.134	0.040	0.080	0.043	0.109	0.035	0.061	0.040	0.070	3060	1.050	0.966	1.097	0.978	1.019	1.100	1.110	1.374	1.571	1.374	1.574
2454	0.367	0.149	0.143	0.035	0.080	0.044	0.107	0.032	0.048	0.031	0.058	3065	1.084	1.234	1.239	1.204	1.109	1.227	1.147	1.834	1.957	1.674	1.849
2455	0.311	0.120	0.143	0.035	0.078	0.043	0.103	0.044	0.035	0.031	0.080	3070	1.088	1.151	1.391	1.484	1.218	1.381	1.256	2.016	1.776	1.885	1.026
2456	0.305	0.117	0.143	0.035	0.074	0.040	0.090	0.044	0.035	0.031	0.080	3075	1.088	1.127	1.135	1.335	1.542	1.263	1.483	2.011	1.775	1.971	1.055
2457	0.362	0.117	0.146	0.037	0.077	0.040	0.112	0.073	0.061	0.095	0.106	3080	1.203	1.249	1.390	1.472	1.250	1.506	1.363	1.961	2.038	1.972	1.254
2470	0.374	0.132	0.129	0.073	0.076	0.046	0.102	0.061	0.083	0.075	0.107	3090	1.120	0.964	1.373	1.344	1.137	1.393	1.279	1.897	1.850	1.921	1.482
2480	0.389	0.121	0.119	0.076	0.082	0.045	0.098	0.067	0.073	0.087	0.104	3095	1.030	0.900	1.399	1.448	1.088	1.272	1.142	2.100	1.775	2.191	1.248
2485	0.393	0.095	0.108	0.037	0.071	0.039	0.093	0.063	0.071	0.069	0.122	3100	1.103	1.037	1.475	1.596	1.293	1.591	1.291	2.429	2.239	2.625	1.490
2490	0.352	0.117	0.137	0.057	0.074	0.044	0.093	0.057	0.084	0.055	0.076	3105	1.091	1.037	1.475	1.596	1.293	1.591	1.291	2.429	2.239	2.625	1.490
2475	0.381	0.138	0.126	0.059	0.079	0.047	0.102	0.061	0.083	0.075	0.107	3110	1.091	1.215	1.466	2.318	1.423	1.951	1.414	2.776	2.726	2.914	1.792
2480	0.389	0.121	0.119	0.076	0.082	0.045	0.098	0.067	0.073	0.087	0.104	3115	1.014	1.286	1.363	2.151	1.383	1.969	1.423	2.629	2.904	2.860	1.801
2485	0.393	0.095	0.108	0.037	0.071	0.039	0.093	0.063	0.071	0.069	0.122	3120	0.892	1.211	1.305	1.915	1.279	1.766	1.344	2.518	2.760	2.714	1.571
2490	0.352	0.117	0.137	0.057	0.074	0.044	0.093	0.057	0.084	0.055	0.076	3125	0.760	1.193	1.310	1.929	1.528	1.626	1.327	2.570	2.947	2.931	1.288
2500	0.456	0.146	0.136	0.044	0.089	0.046	0.123	0.059	0.060	0.076	0.080	3130	0.892	1.211	1.305	1.915	1.279	1.766	1.344	2.518	2.760	2.714	1.571
2505	0.459	0.131	0.144	0.052	0.093	0.048	0.126	0.060	0.055	0.083	0.092	3135	0.760	1.193	1.310	1.929	1.528	1.626	1.327	2.570	2.947	2.931	1.288
2510	0.406	0.135	0.135	0.057	0.077	0.046	0.116	0.067	0.063	0.068	0.092	3140	0.892	1.211	1.305	1.915	1.279	1.766	1.344	2.518	2.760	2.714	1.571
2515	0.317	0.096	0.135	0.057	0.077	0.046	0.116	0.067	0.063	0.068	0.092	3145	1.029	1.233	1.432	2.063	1.323	1.866	1.444	2.948	3.267	3.393	0.883
2520	0.264	0.100	0.112	0.040	0.067	0.041	0.103	0.051	0.053	0.059	0.066	3150	1.029	1.268	1.432	2.078	1.323	1.854	1.484	2.993	3.748	3.471	1.040
2525	0.264	0.100	0.112	0.040	0.067	0.041	0.103	0.051	0.053	0.059	0.066	3155	0.954	1.266	1.219	2.061	1.218	1.766	1.478	2.898	3.863	3.677	1.742
2530	0.268	0.108	0.119	0.050	0.076	0.054	0.116	0.077	0.045	0.084	0.098	3160	0.938	1.123	1.089	2.014	1.146	1.658	1.402	2.642	3.893	3.742	1.588
2535	0.300	0.127	0.130	0.056	0.089	0.072	0.127	0.091	0.067	0.087	0.132	3165	0.792	1.289	1.526	2.344	1.273	2.014	1.436	3.393	4.537	4.637	1.530
2540	0.360	0.142	0.136	0.059	0.103	0.091	0.129	0.091	0.091	0.081	0.148	3170	0.877	1.368	1.294	2.713	1.316	2.087	1.403	3.924	4.137	4.416	1.598
2545	0.360	0.142	0.136	0.059	0.103	0.091	0.129	0.091	0.091	0.081	0.148	3175	0.888	1.430	1.270	2.553	1.432	2.283	1.624	4.556	4.958	4.958	1.213
2550	0.419	0.196	0.164	0.079	0.130	0.084	0.147	0.093	0.103	0.089	0.149	3180	0.888	1.430	1.270	2.553	1.432	2.283	1.624	4.556	4.958	4.958	1.213
2555	0.409	0.209	0.177	0.079	0.130	0.084	0.147	0.093	0.103	0.089	0.149	3185	0.824	1.314	1.211	2.312	1.363	2.118	1.531	3.527	4.571	5.118	1.395
2560	0.429	0.215	0.200	0.084	0.143	0.079	0.170	0.094	0.096	0.112	0.138	3190	0.806	1.216	2.239	1.333	2.003	1.452	3.378	4.580	4.940	1.676	
2565	0.478	0.298	0.232	0.084	0.162	0.086	0.189	0.102	0.104	0.099	0.154	3195	0.792	1.289	1.526	2.344	1.273	2.014	1.436	3.393	4.537	4.637	1.530
2570	0.521	0.342	0.207	0.085	0.132	0.082	0.191	0.071	0.074	0.044	0.070	3200	0.765	1.397	1.642	2.493	1.286	2.015	1.509	3.924	4.137	4.416	1.598
2575	0.521	0.342	0.207	0.085	0.132	0.082	0.191	0.071	0.074	0.044	0.070	3205	0.498	1.455	1.398	2.487	1.289	2.161	1.517	3.924	4.137	4.416	1.598
2580	0.507	0.324	0.207	0.071	0.140	0.089	0.182	0.082	0.090	0.044	0.073	3210	0.654	1.441	1.207	2.286	1.219	2.078	1.440	3.100	3.860	4.266	2.031
2585	0.507	0.324	0.207	0.071	0.140	0.089	0.182	0.082	0.090	0.044	0.073	3215	0.617	1.332	1.051	2.104	1.176	1.928	1.330	2.782	3.569	4.204	1.484
2590	0.461	0.237	0.226	0.075	0.158	0.098	0.169	0.096	0.074	0.153	0.127	3220	0.659	1.305	1.051	1.946	1.064	1.916	1.287	2.298	3.658	4.322	1.661
2600	0.405	0.173	0.215	0.107	0.149	0.101	0.159	0.109	0.142	0.158	0.121	3225	0.690	1.314	1.276	1.870	1.031	1.848	1.269	2.202	4.043	4.348	1.731
2605	0.363	0.213	0.209	0.121	0.153	0.101	0.174	0.131	0.232	0.177	0.079	3230	0.692	1.051	1.436	1.874	0.953	1.718	1.144	2.108	4.038	4.348	1.731
2610	0.336	0.197	0.216	0.110	0.156	0.101	0.178	0.131	0.232	0.177	0.079	3235	0.663	1.102	1.394	2.082	1.207	1.859	1.217	2.108	4.038	4.348	1.731
2615	0.334	0.204	0.228	0.116	0.165	0.104	0.184	0.133	0.246	0.192	0.146	3240	0.661	1.102	1.394	2.082	1.207	1.859	1.217	2.108	4.038	4.348	1.731