SOLAR BURST WITH MILLIMETRE-WAVE EMISSION AT HIGH FREQUENCY

P.Kaufmann, E. Correia, J.E.R. Costa, A.M. Zodi Vaz

Instituto de Pesqui**s**as Espaciais, CNPq Brasil

and

B. R. Dennis

NASA, Goddard Space Flight Center, USA

ABSTRACT. Solar burst emission data at shorter millimetre wavelengths have been based on a few events measured with low sensitivity and relatively poor time resolution. We present here the first high sensitivity and high time-resolution observations taken simultaneously at 90 GHz (λ = 3.3 mm) and at 30 GHz (λ = 10 mm). These have identified a unique impulse burst on 21 May 1984 with fast pulsed emission that was considerably more intense at 90 GHz than at lower frequencies. Hard X-ray time structures at energies above 25 keV were almost identical to the 90 GHz structure to better than 1s. The structure of the onset of the major 90-GHz burst coincided with the hard X-ray structure to within 128 ms. All 90-GHz major time structures consisted of trains of multiple sub-second pulses with rise times as short as 0.03s and amplitudes that were large compared with the mean flux. When detectable, the 30-GHz sub-second pulses had smaller relative amplitude and were in phase with the corresponding 90-GHz pulses.

(Published in Nature, 31 January, 1985).

E. Correia, J.E.R. Costa, P. Kaufmann, and A.M. Zodi Vaz: INPE, Instituto de Pesquisas Espacia: CNPq, Caixa Postal 515, 12.200 São José dos Campos, SP Brasil.

B.R. Dennis: NASA, Goddard Space Flight Center, Laboratory for Astronomy and Solar Physics, Greenbelt, Maryland 20771, USA.