IONIZATION CORRECTION FACTORS IN PLANETARY NEBULAE: NITROGEN

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We have recently derived a set of new ionization correction factors (ICFs) to be used in planetary nebulae. Here we present the results for nitrogen. Significant differences are found between the N/O values derived from our ICF and the ones calculated through the usual method: N/O = N⁺/O⁺.

Ionization correction factors (ICFs) are essential to estimate total abundances in PNe. Using a large grid of photoionization models, computed with Cloudy c10.00 (Ferland et al. 1998), we derived analytical expressions for the ICFs (and for the error bars associated with the ICFs) of He, O, N, Ne, S, Ar, Cl, and C. The whole analysis will be presented elsewhere (Delgado-Inglada, Morisset, & Stasinska, 2014, in prep.)

The upper panel of Figure 1 shows the values of $ICF_m(N^+/O^+) = x(O^+)/x(N^+)$ as a function of $O^{++}/(O^++O^{++})$ for the models. We see from the figure that the N/O values calculated as N/O = N^+/O^+ (dashed line) will be overestimated (underestimated) in PNe with high (low) values of He⁺⁺/(He⁺+He⁺⁺). We suggest a new ICF_f(N⁺/O⁺) based on O⁺⁺/(O⁺+O⁺⁺) and He⁺⁺/(He⁺+He⁺⁺). Each of the four solid lines in the plot represents our ICF for one specific value of He⁺⁺/(He⁺+He⁺⁺).

In the lower panel of this figure we show the uncertainties associated with our ICF (solid lines) and with the usual correction scheme for nitrogen, N/O = N⁺/O⁺, proposed by Kingsburgh & Barlow (1994). The relative differences between ICF_m(N⁺/O⁺) and ICF_f(N⁺/O⁺) are displayed as a function of O⁺⁺/(O⁺+O⁺⁺). The uncertainties associated with our ICF are significantly lower than those associated with the one adopted in Kingsburgh & Barlow (1994). The differences are more evident in the models with low He⁺⁺/(He⁺+He⁺⁺) values (corresponding to low effective temperature central



Fig. 1. Upper panel: Values of $\rm ICF_m(N^+/O^+)$ as a function of $\rm O^{++}/(O^++O^{++})$. Lower panel: Values of $\rm (ICF_m(N^+/O^+)\text{-}ICF_f(N^+/O^+))/ICF_f(N^+/O^+)$ as a function of $\rm O^{++}/(O^++O^{++})$.

stars), where N/O could be underestimated by up to 0.4 dex when calculated simply as N⁺/O⁺. The uncertainties in log(N/O) associated with our ICF are in general lower than ± 0.2 dex.

REFERENCES

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