THE ZADKO TELESCOPE RESULTS: TEN YEARS OF SCIENCE

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RESUMEN

El telescopio Zadko de 1,0 m y focal f/4 de giro rápido se encuentra en Australia Occidental, aproximadamente setenta kilómetros al norte de Perth en Yeal, en la Comarca de Gingin, en un área dedicada de "baja luminosidad". Es la única instalación de investigación óptica en su clase en esta latitud sur entre la costa este de Australia y Sudáfrica y puede captar rápidamente imágenes de transitorios ópticos en una longitud no monitoreada por otras instalaciones similares. Revisamos aquí los principales resultados alcanzados durante la última década y damos algunas pistas hacia las metas establecidas para los próximos años. Finalmente discutimos las modificaciones y mejoras que tuvimos que realizar en las instalaciones para alcanzar estos nuevos objetivos.

ABSTRACT

The 1.0 meter f/4 fast-slew Zadko telescope is located in Western Australia, approximately seventy kilometers north of Perth at Yeal in the Shire of Gingin in a dedicated "low-luminosity" area. It is the only meter class optical research facility at this southern latitude between the east coast of Australia and South Africa and can rapidly image optical transients at a longitude not monitored by other similar facilities. We review here the main results achieved during the last decade and give some points toward the goals set for future years. Finally we discuss the modifications and improvements we had to perform in the facility to reach these new goals.

Key Words: instrumentation: detectors — telescopes

1. INTRODUCTION

The Zadko Telescope (Coward et al. 2010) is a 1.0 meter f/4 Ritchey-Chretien telescope located in Western Australia, in the shire of Gingin. It has been constructed by DFM Engineering Ltd, and donated to the University of Western Australia by resource company Claire Energy CEO, J. Zadko. The telescope has been in operation since 1^{st} April 2009. The main properties of the observatory are listed in Table 1.

In mid 2010, the telescope control system was fully upgraded using the TAROT robotic telescope software system (Klotz et al. 2008). This suite of independently running programs comprises two main components: AudeLA⁵ and ROS (Robotic Observatory Software). There is also a third party software (mostly drivers) for the mechanical and ancillary systems of the observatory (i.e. the telescope and the weather station). All technical details are reported in Moore et al. (2021).

TABLE 1

SPECIFICATIONS OF THE OBSERVATORY		
	Property	Value
	Observatory code	D20
	Longitude	115° 42' 49" E
	Latitude	$31^{\circ} 21' 24'' S$
	Altitude	50 meters
	Seeing	Variable $(1-4'')$
	Limiting magnitude	$R{\sim}21.5$ (180s exposure)
	Camera	QHY 160M
	Filters	Sloan g', r', i', Clear

The key feature of the telescope is the optimization for performing automated follow-up from alerts or triggers from different external instruments. In addition, it employs a dynamic scheduling system that allows other projects to operate when not in alert mode. This sets the facility apart from many other systems that are "internet" telescopes that still require some level of manual monitoring. Furthermore the location of the facility at a longitude not covered by many other meter class facilities provides an important resource for many projects.

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⁵http://www.audela.org/english_audela.php

125

The scientific goal of the Zadko telescope is to explore a previously uncharted region of the 'transient sky' which includes:

- Gamma ray bursts: Rapid optical follow-up of *Swift* Alerts.
- Antares neutrino alert follow-up.
- Solar system studies: follow-up of asteroids and studies of "Barbarian asteroids".
- Gravitational wave follow-up of events detected by LIGO-Virgo.
- Nearby supernovae and lensed supernovae search.
- Education outreach and training.
- Emerging projects including automated followup of radio transients and space debris tracking.

The Zadko Telescope having fulfilled its purpose over a ten year period and producing many publications was shut down in late 2018 for a long overdue maintenance. This included shipping both Primary and Secondary mirrors back to the USA for stripping and re-aluminization (see Moore et al. 2021, for details).

In this paper, we briefly review the results obtained with the telescope on gamma-ray bursts (section 2), and gravitational waves (Section 3) taken before the shut down, and discuss the future of the instrumentation (Section 4).

2. GAMMA-RAY BURSTS

We are using a similar strategy to the TAROT network to perform the follow-up of GRBs detected by high energy instruments. In short, the telescope first performs a trailed observation for 30 seconds, allowing a temporal resolution of about 1 second of the prompt light curve, continuing with observations using the clear filter for the next hour. After that time, where the temporal coverage of the afterglow can allow for discontinuous observations, we are alternating filters to gather color information. The observations are repeated the next night in case of bright events.

Most of the results have been reported in Coward et al. (2017), and only two events have not been published yet: GRB 160127A (Crisp et al., in preparation), and GRB 170202A (Gendre et al., in preparation). During nine years of activity the Zadko telescope performed 39 follow-ups of GRBs and detected 13 of them. This rate is common for a robotic telescope placed in a remote location.

3. GRAVITATIONAL WAVES

The scientific interest of the telescope has slightly shifted from gamma-ray bursts to the counterparts of gravitational wave emitters with the detections of the first events by the LIGO-Virgo collaboration (Abbott et al. 2016, 2017). The Zadko telescope has been part of the follow-up campaign of GW170817, and has provided very important optical data (Andreoni et al. 2017). Its limited field of view, however, prevents us to realistically find the counterpart inside a large position error box. The telescope entered the GRANDMA collaboration (Antier et al. 2020) for that purpose. The GRANDMA collaboration aims at coordinating observations between telescopes with a large field of view (to detect potential counterparts) and ones with a small field of view (to confirm/reject the association). Since the start of the run O3, the Zadko telescope participated in three observation campaigns of LIGO-Virgo events. These observations are reported in Antier et al. (2020).

4. FUTURE OF THE INSTRUMENTATION

Maintenance of the telescope has allowed for a pause in the observations and provided time to reflect on the scientific goals. In addition, the experience gained during the first follow-up of fast radiobursts and gravitational wave events has shown that an evolution of the telescope instrumentation was required. Initially we replaced the previous camera (being an Andor Ikon-L with a CCD) with a new QHY camera equipped with a CMOS chip. We chose a low price camera with a small field of view in order to test this new technology and to be sure it would face the challenges of our scientific goals. Having now confirmed that this technology is suitable for our needs, we are seeking to adapt a large CMOS camera behind the telescope.

We have also confirmed that we require a larger choice of filters to fulfill all our requirements, thus modifying the whole optical path due to limitations with our existing filter wheel. This led us to redesign the mechanical link between the instrument and the telescope enabling an increase in the number of instruments one could mount on the telescope. Lastly, we are investigating the possibility to have a full field low resolution spectrometer, which would allow us to discriminate the various kinds of transients when they are detected, and not during a later follow-up.

It is expected that with all these new modifications, and the mechanical refurbishment of the telescope discussed in Moore et al. (2021), the Zadko telescope will be able to fulfill its scientific goals for another ten fruitful years.

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