3D Models of Stellar Interactions

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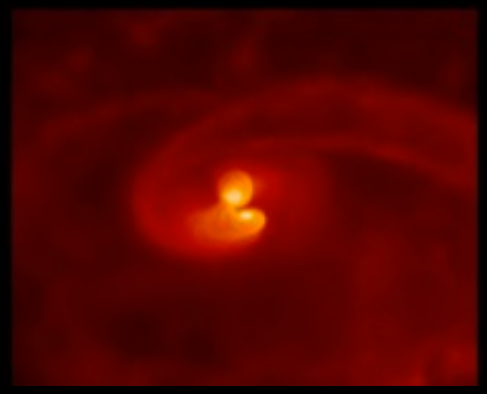
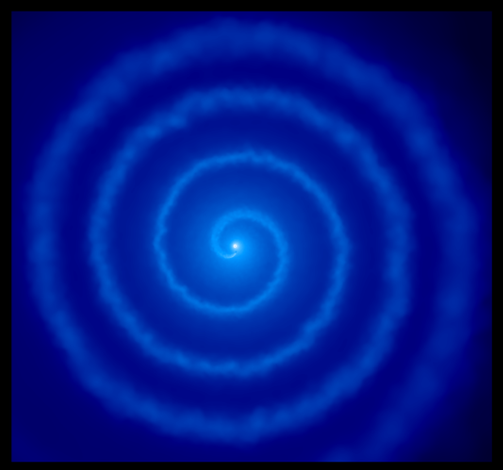
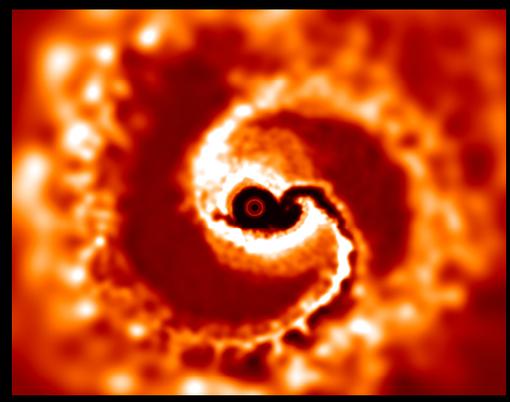
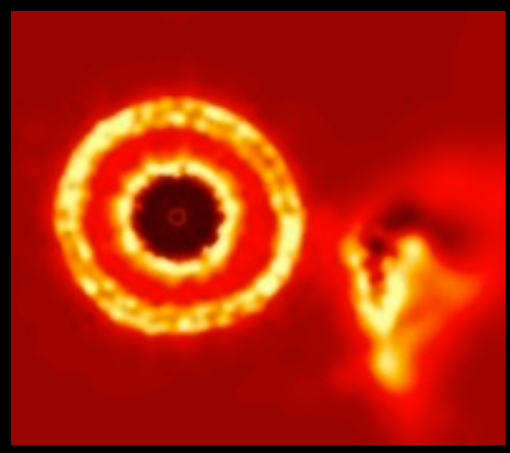


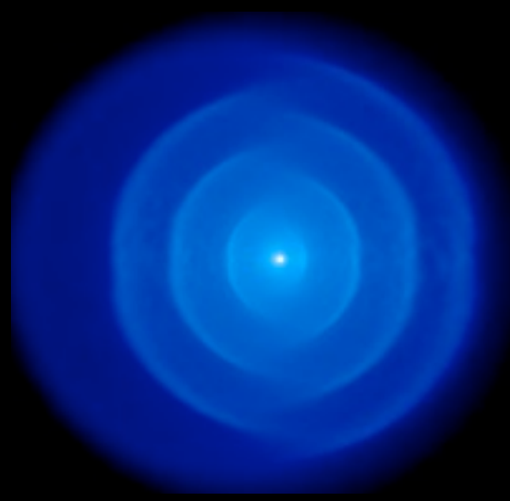


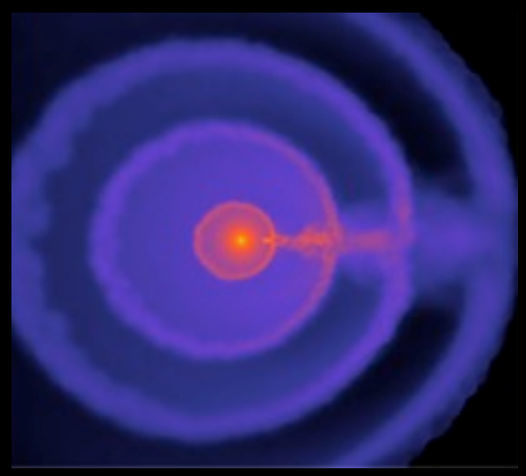
Stellar wind interactions

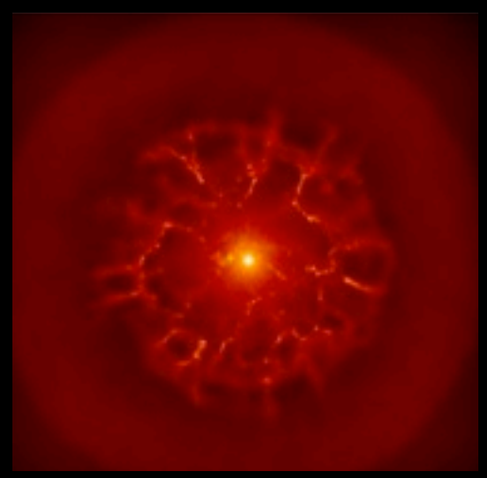
- wind-ISM interactions
- wind-wind interactions
- binary wind interactions

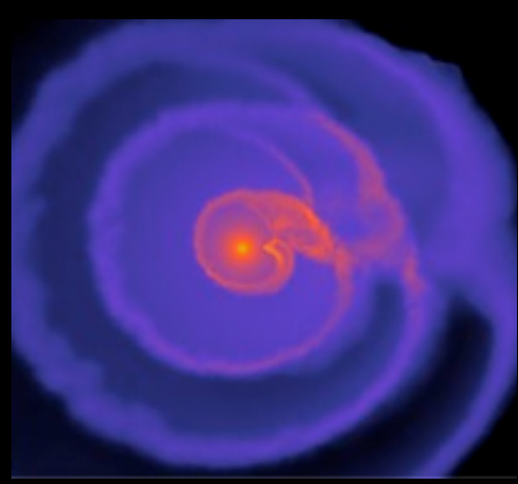


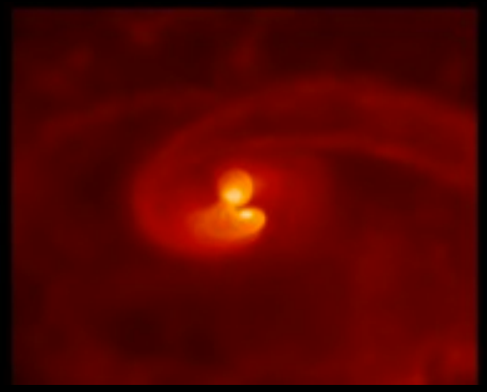


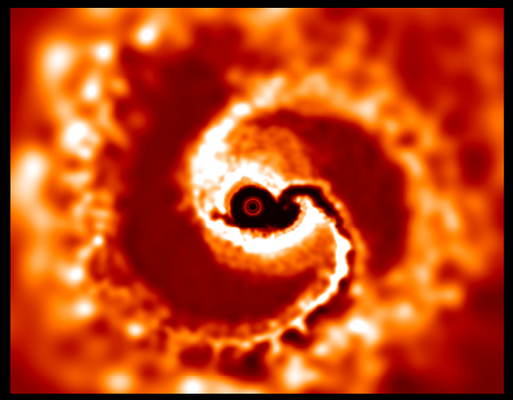






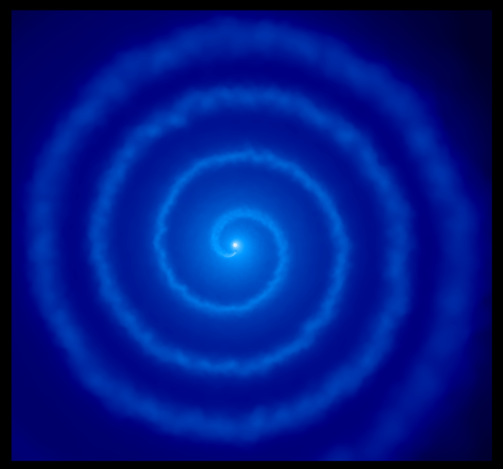


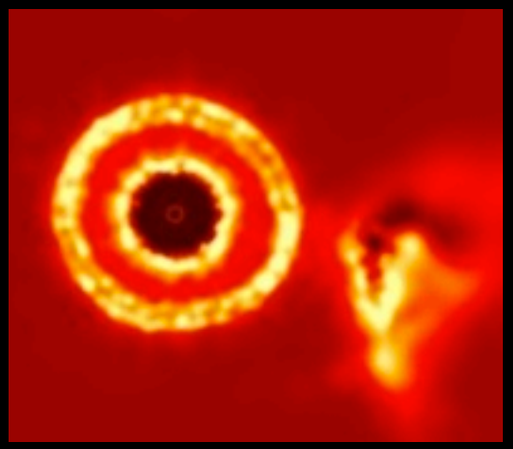




Summary

- Wind interactions can produce a wide range of envelope morphologies
  - spirals
  - arcs
  - shells
  - equatorial outflows





- WRLOF: focusing can explain density enhancement needed for bi-polars, e.g. M2-9
- Mass transfer rates are higher than BHL wind accretion (cf. powering jets)