

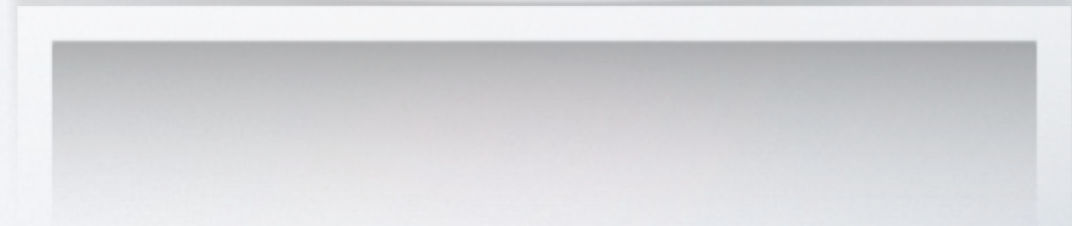
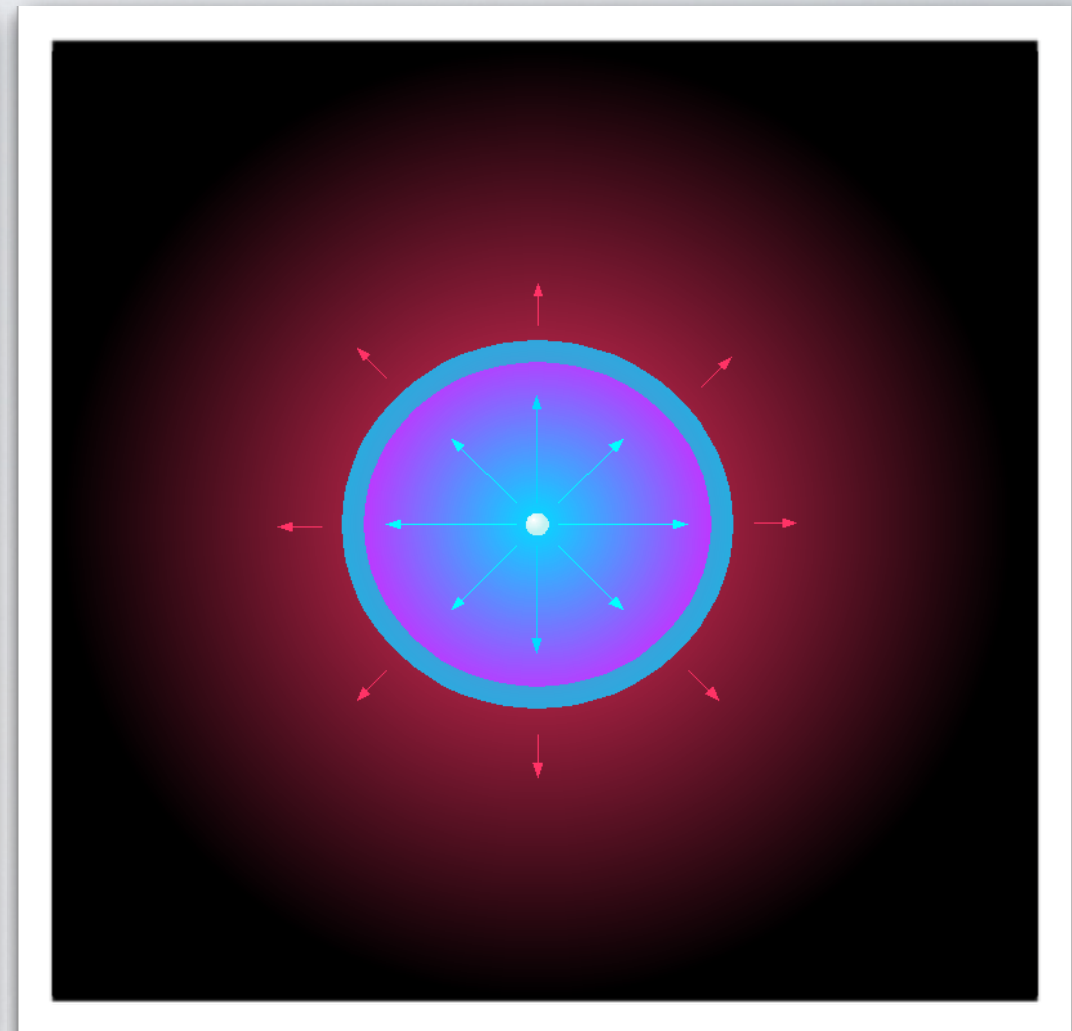
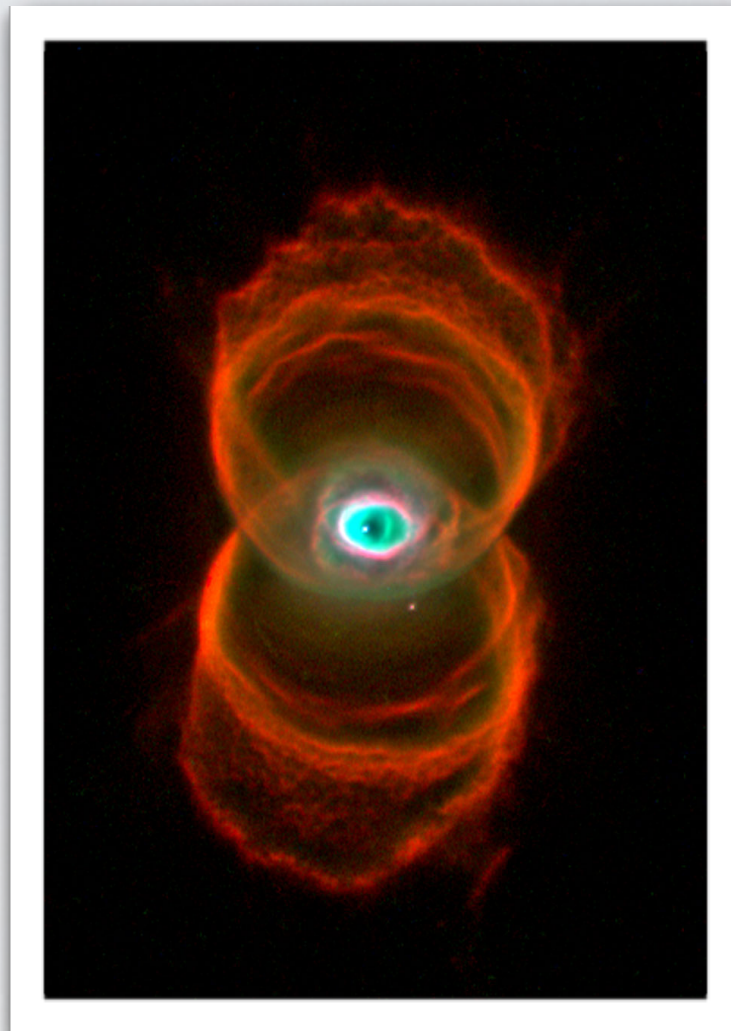
THE MORPHO-KINEMATICS OF PLANETARY NEBULAE WITH BINARY CENTRAL STARS

Miguel Santander-García
David Jones

I. HISTORY - ASPHERICAL PLANETARY NEBULAE

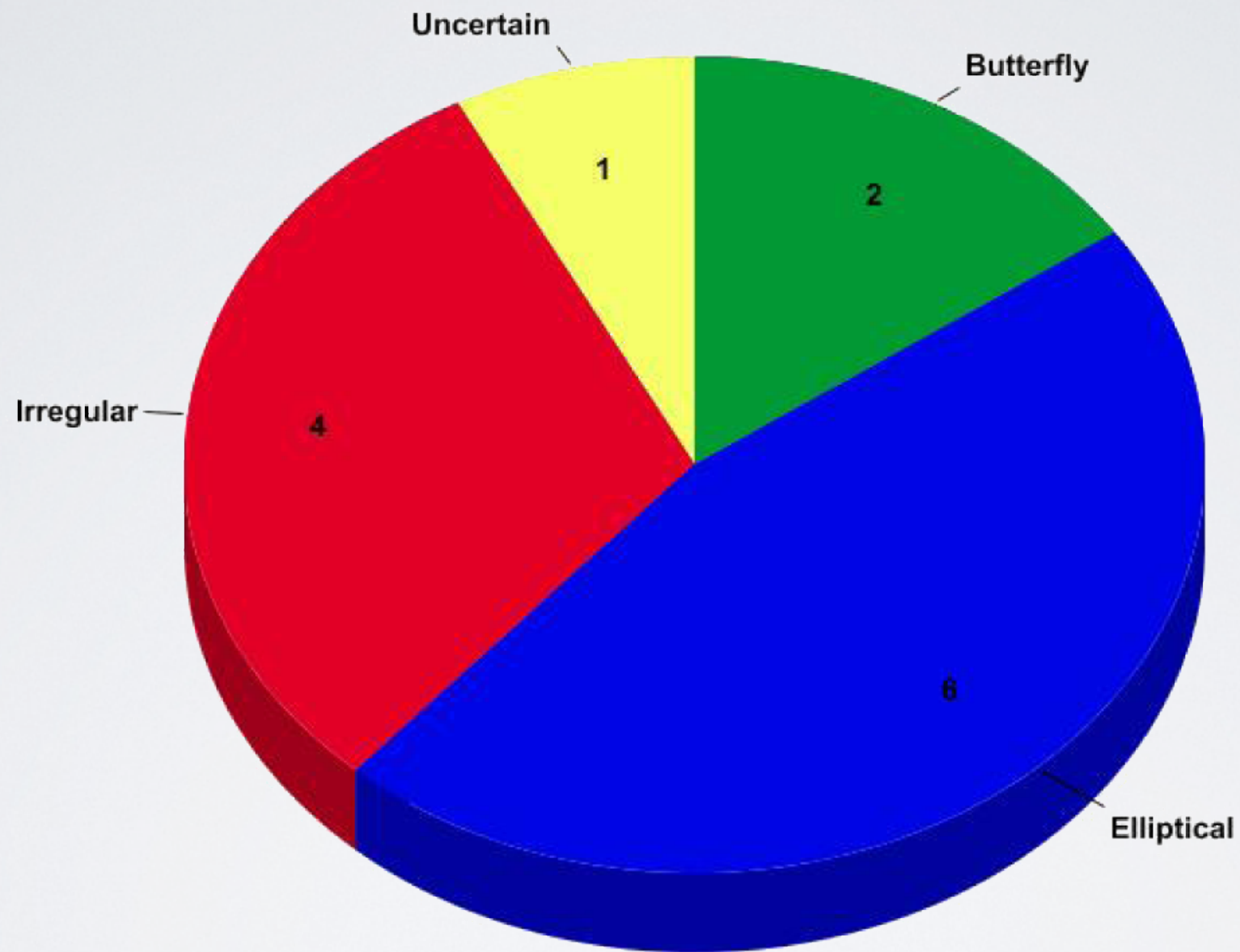
e.g. De Marco 2009, PASP, 121, 316

- Stellar rotation
- Magnetic fields
- Binaries!



I. HISTORY - THE BEGINNING

Bond & Livio 1990, ApJ, 355, 568



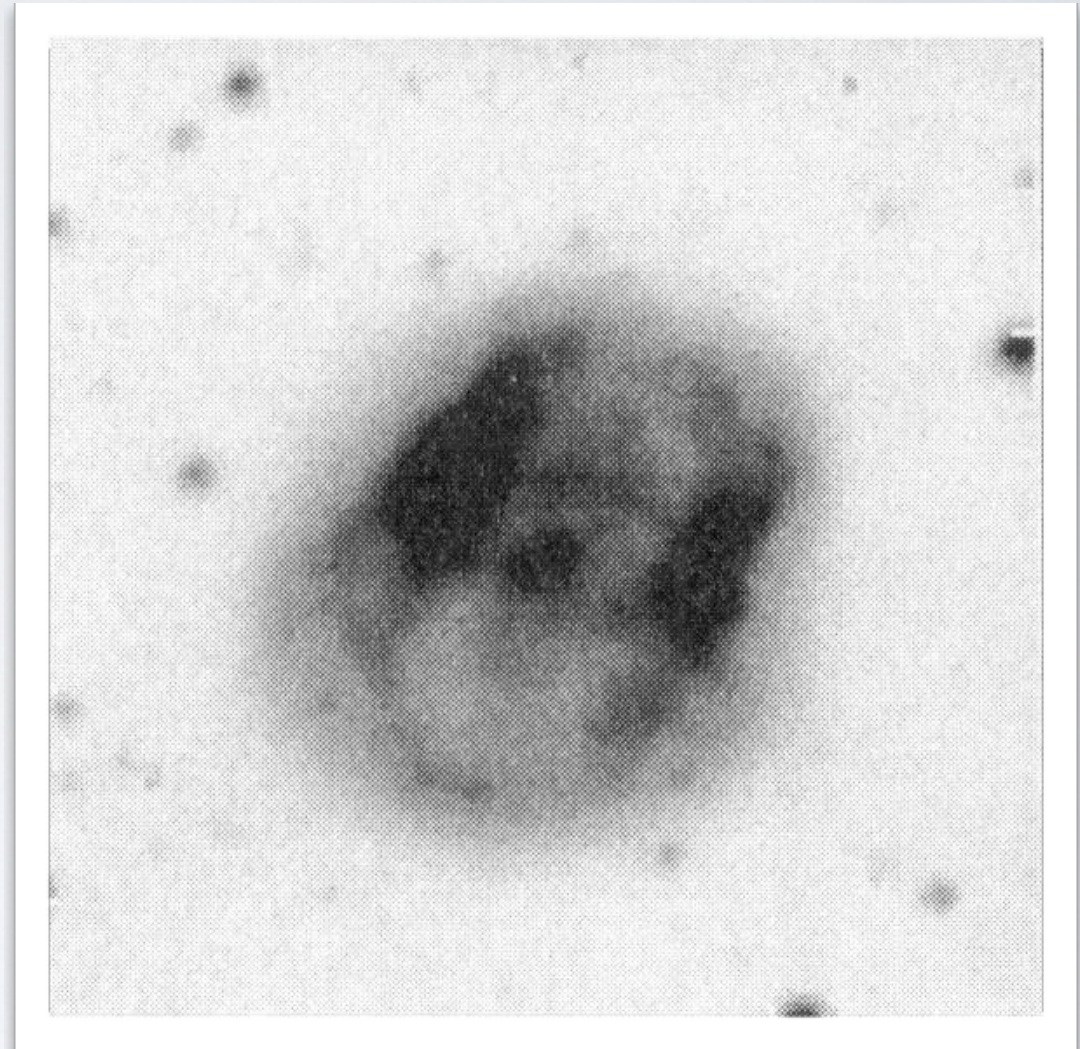
The problem: *Fraction of non-spherical PNe ($\approx 80\%$) much larger than fraction known to contain close-binary nuclei ($\sim 15\%$)*

I. HISTORY - DEEPER IMAGERY

Pollacco & Bell 1997, MNRAS, 284, 32

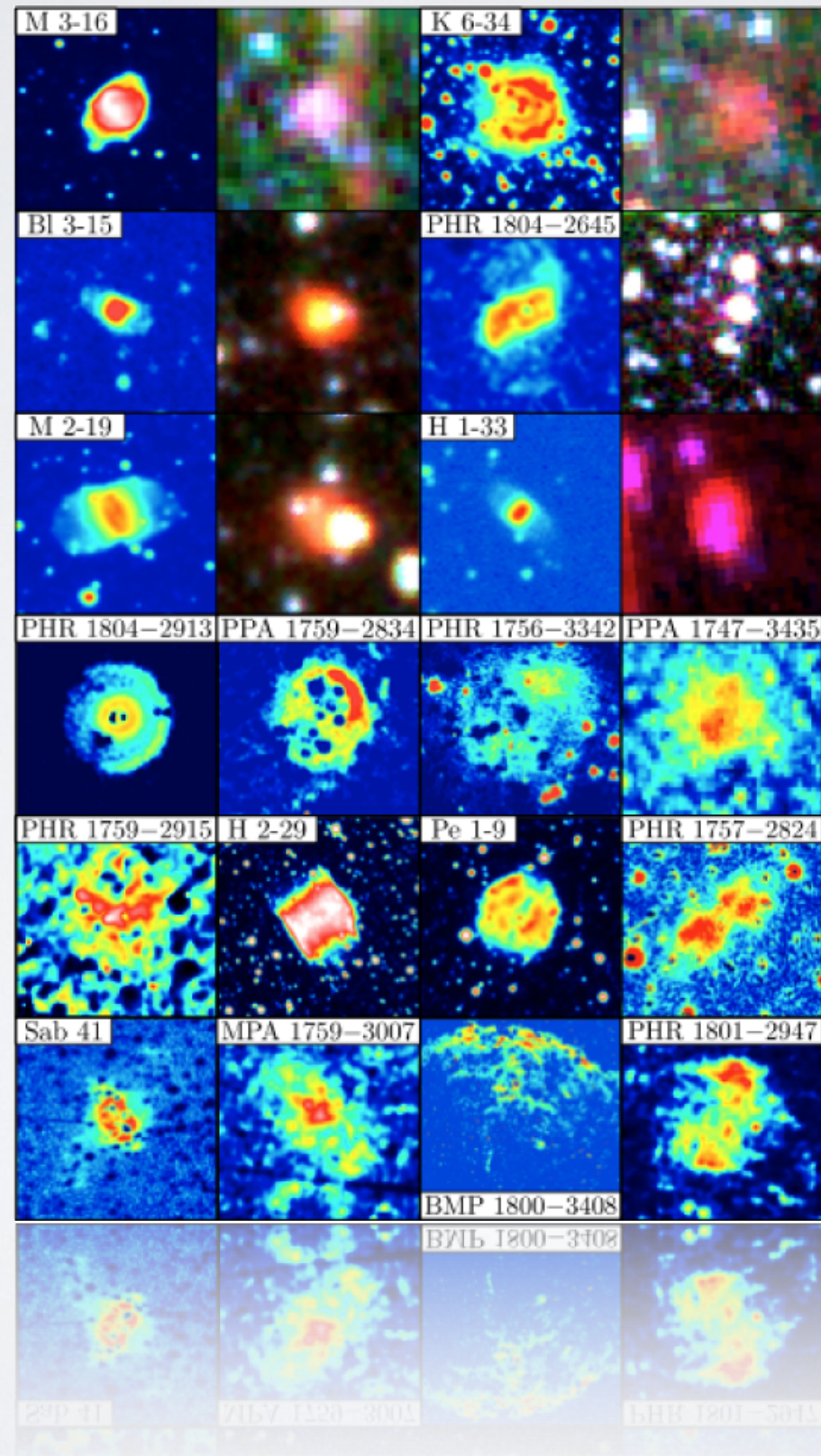
Abell 41

- Deep narrowband $H\alpha + [NII]$ imagery reveals extremely interesting features
- Only 4 PNe but all bipolar/elliptical
- 3/3 apparent inclination matches (nebular axis-orbital plane)



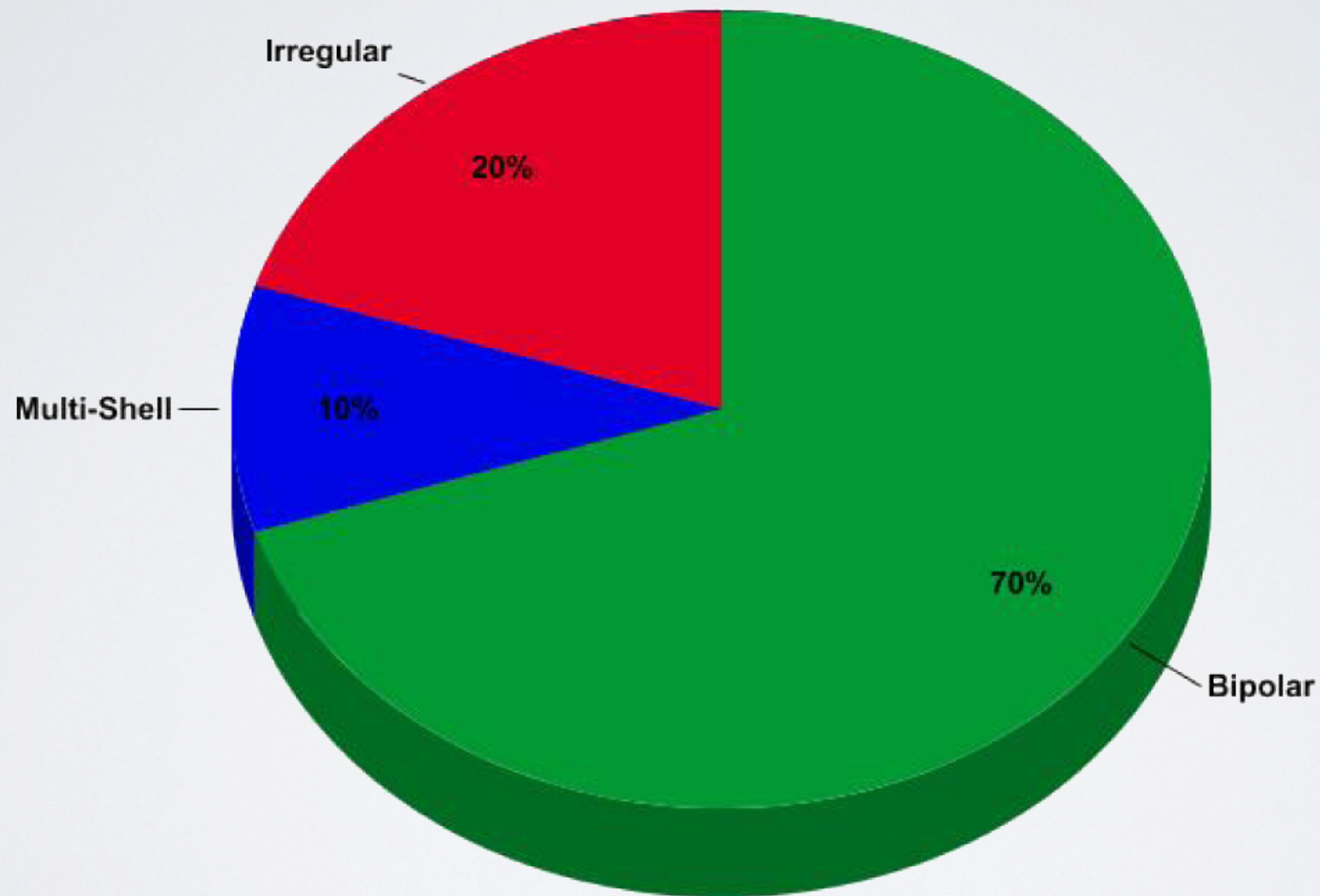
I. HISTORY - BIGGER SAMPLE

Miszalski et al. 2009, A&A, 505, 249



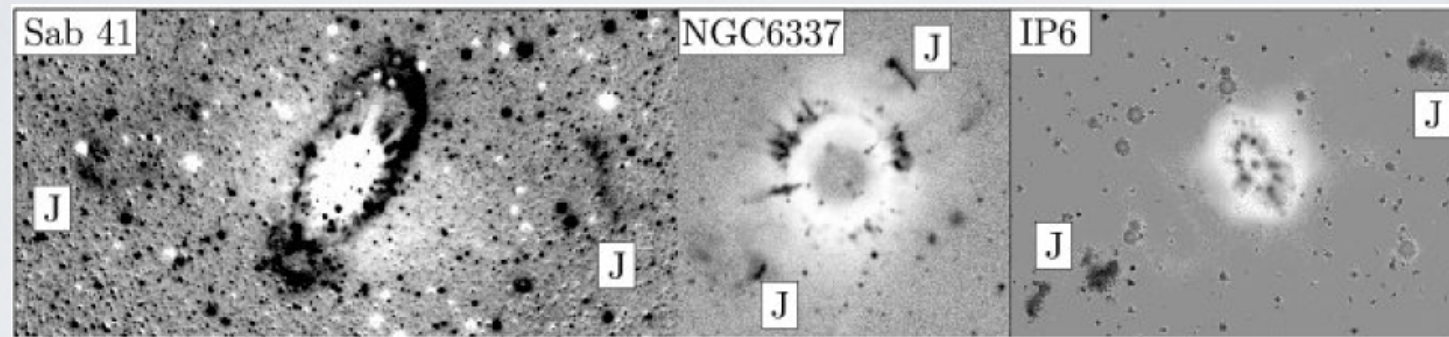
I. HISTORY - BIGGER SAMPLE

Miszalski et al. 2009, A&A, 505, 249



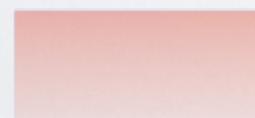
2. BINARY INDICATORS?

Miszalski et al. 2009, A&A, 505, 249



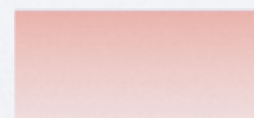
LIS

LIS



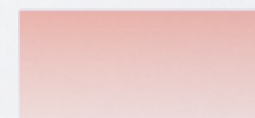
Rings

Rings



Jets

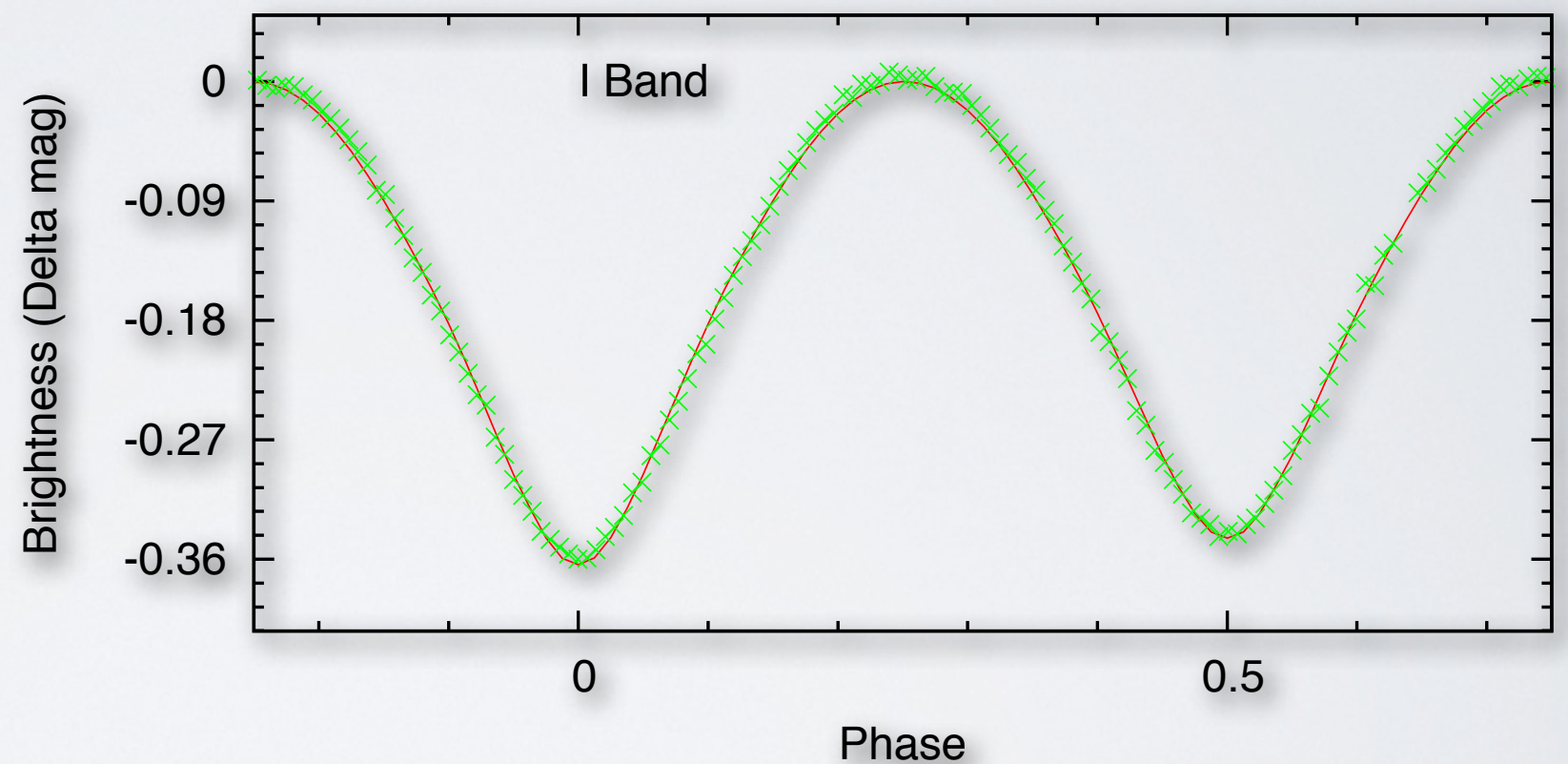
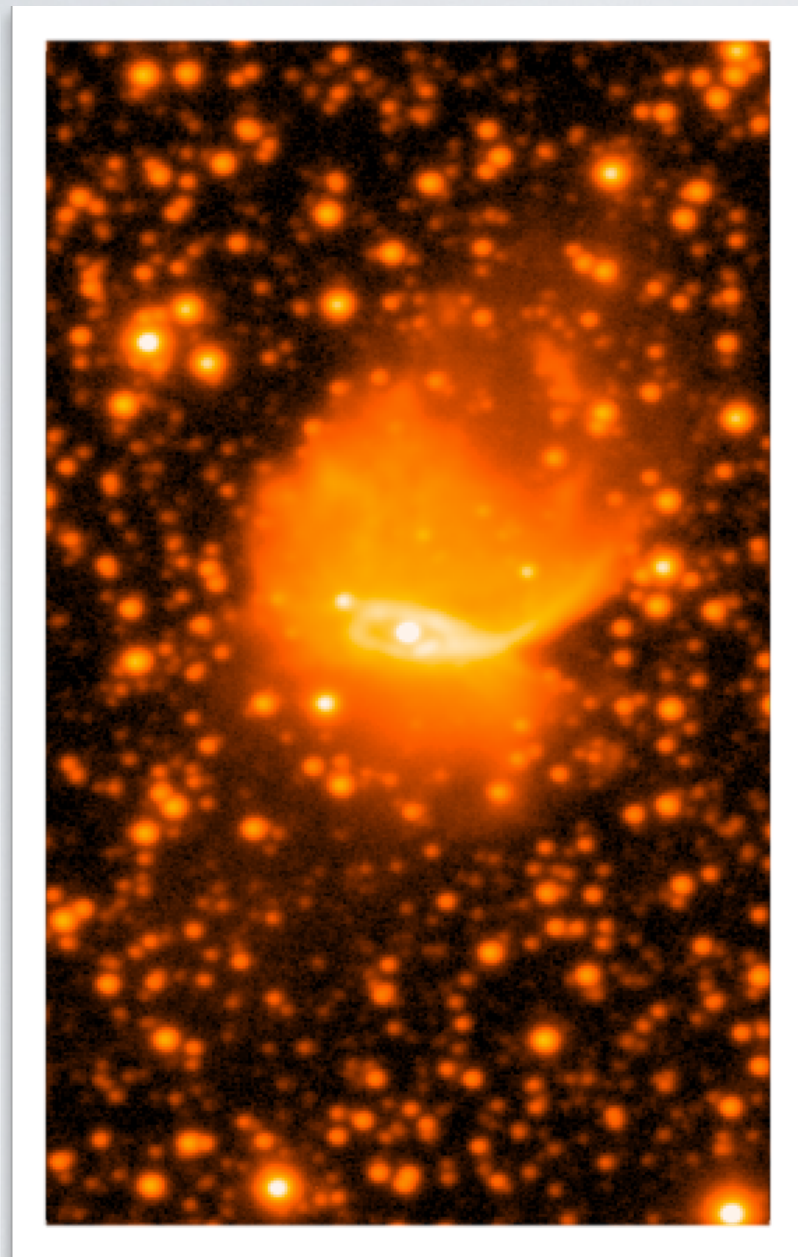
Jets



2. BINARY INDICATORS!

Santander-García et al., in preparation

Hen 2-428

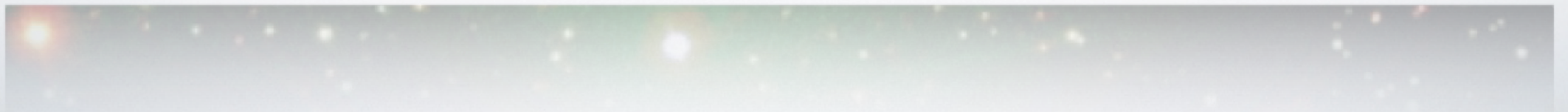


$P = 4.2$ hours

2. BINARY INDICATORS!

Boffin et al. 2012, Science, 338, 773

Fg I



SOMETHING ABOUT THE MORPHOLOGIES...

- Bipolars prevail
- No round nebulae. Irregulars/multiple shells?
- Imagery can be highly ambiguous
- Limited information from images alone...

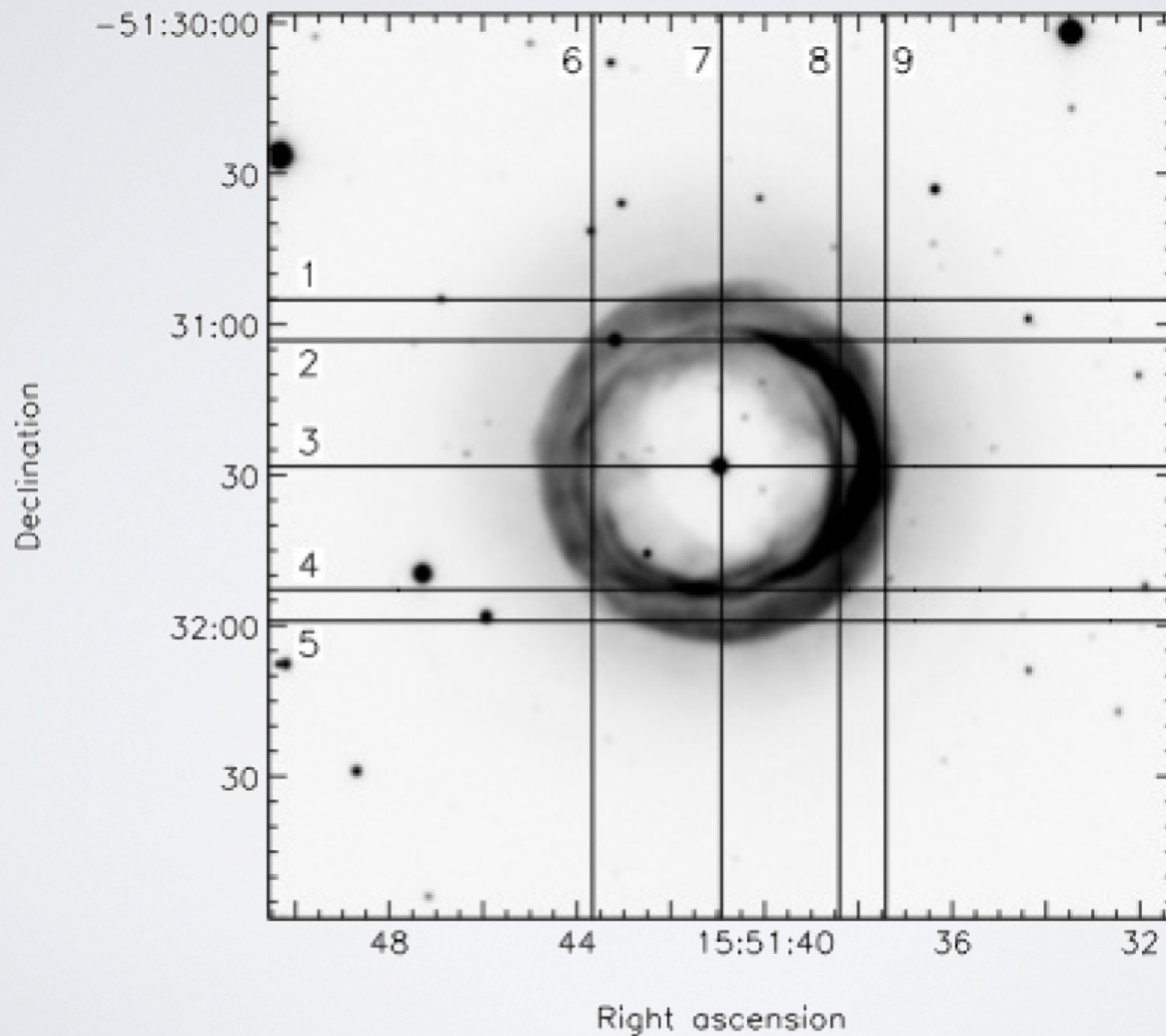
3. KINEMATICS - AN EXAMPLE

Sp I



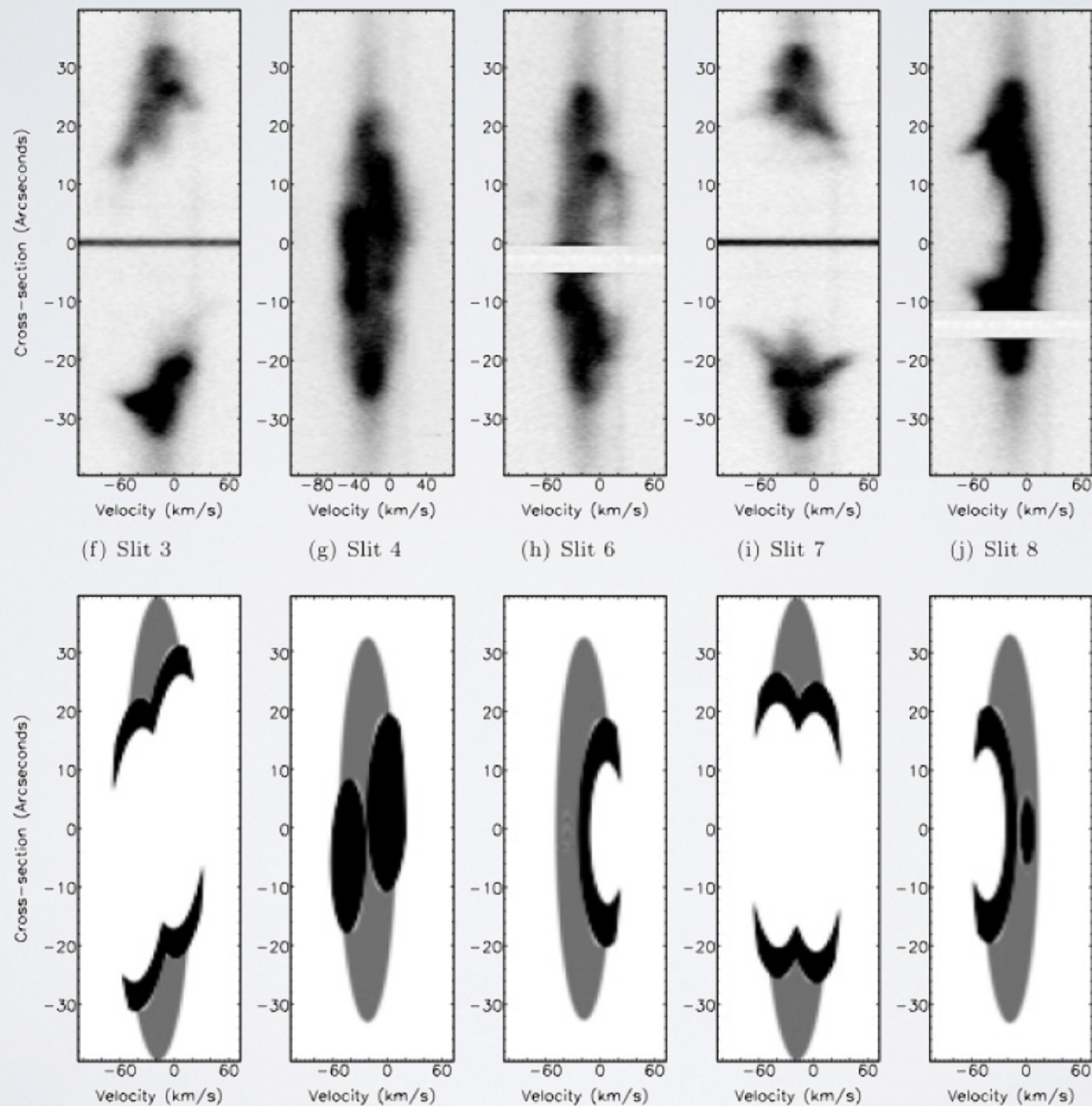
3. KINEMATICS - AN IMAGE MIGHT BE WORTH 1000 WORDS...

Jones et al. 2012, MNRAS, 420, 2271



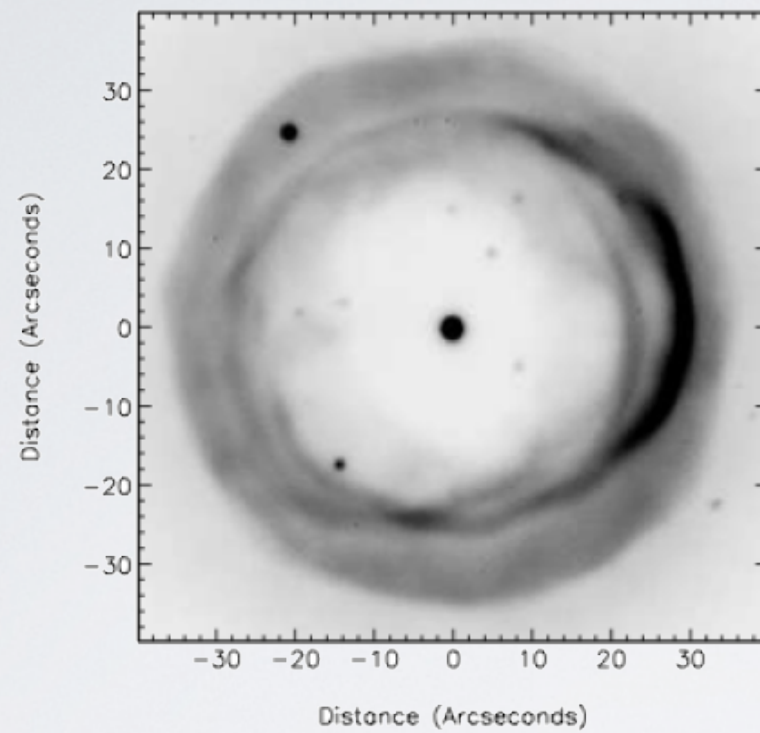
3. KINEMATICS - ...BUT A SPECTRUM IS WORTH 1000 IMAGES!

Jones et al. 2012, MNRAS, 420, 2271

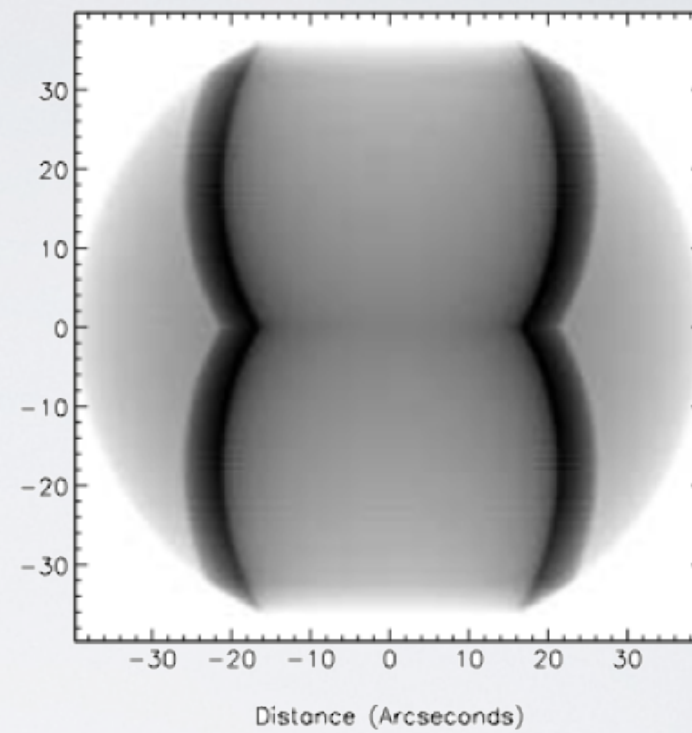


3. KINEMATICS - NOW IN 3D!

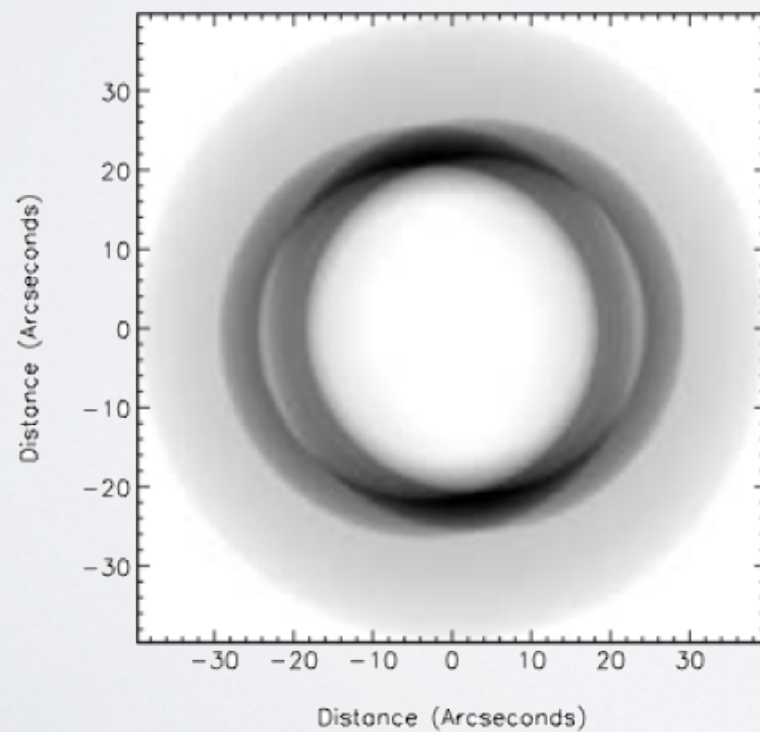
Jones et al. 2012, MNRAS, 420, 2271



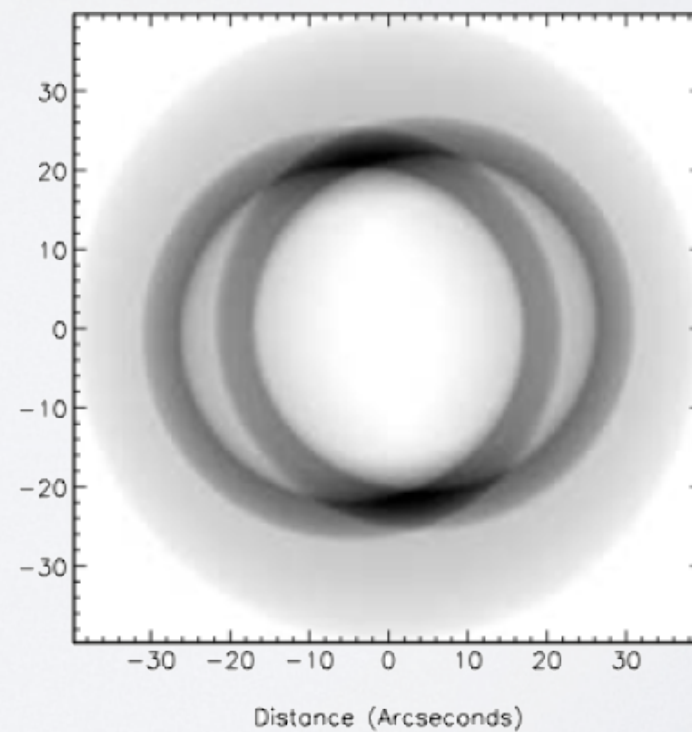
(a) [O III] $\lambda 5007$ Å



(b) 90°



(c) 10°

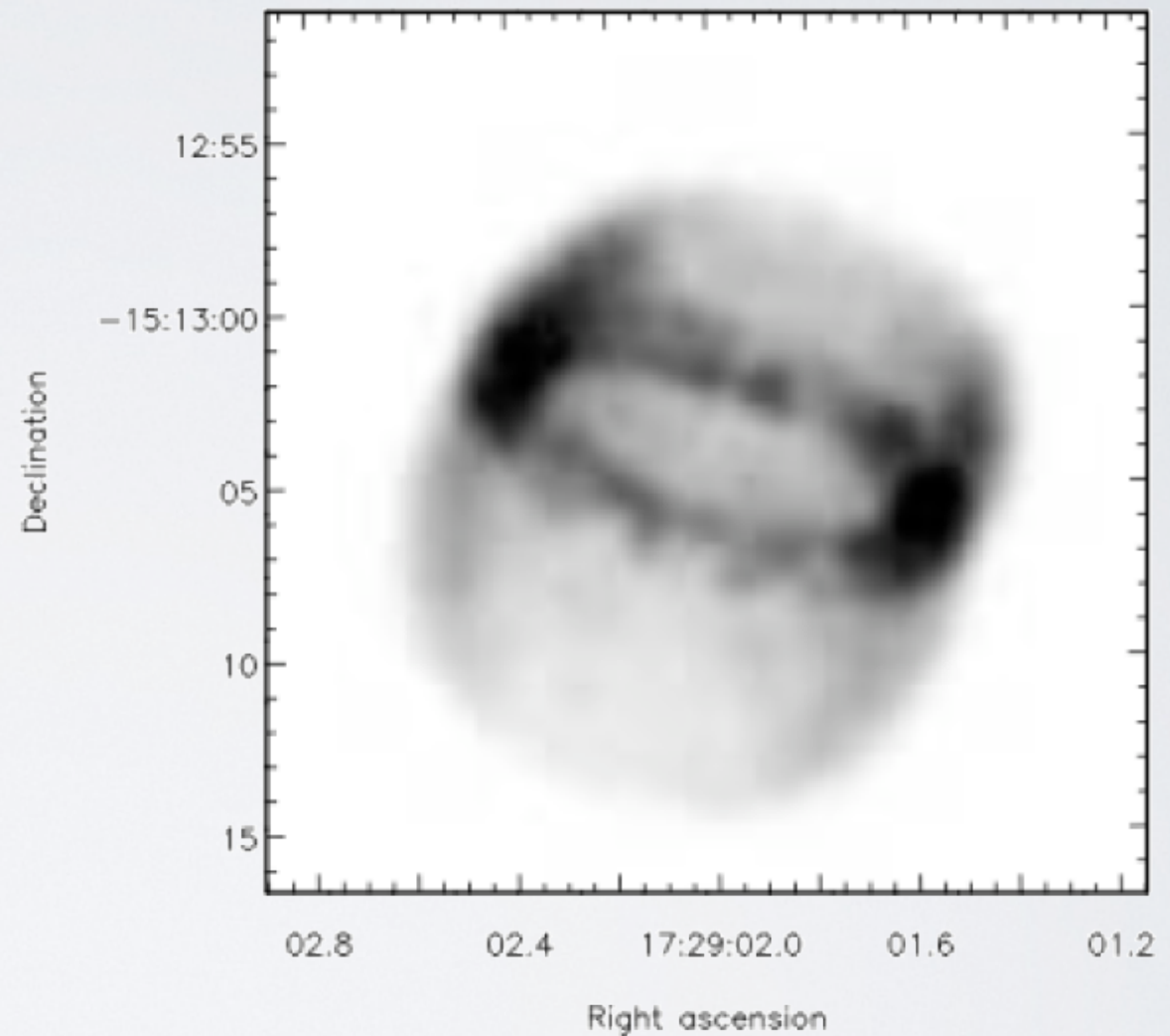
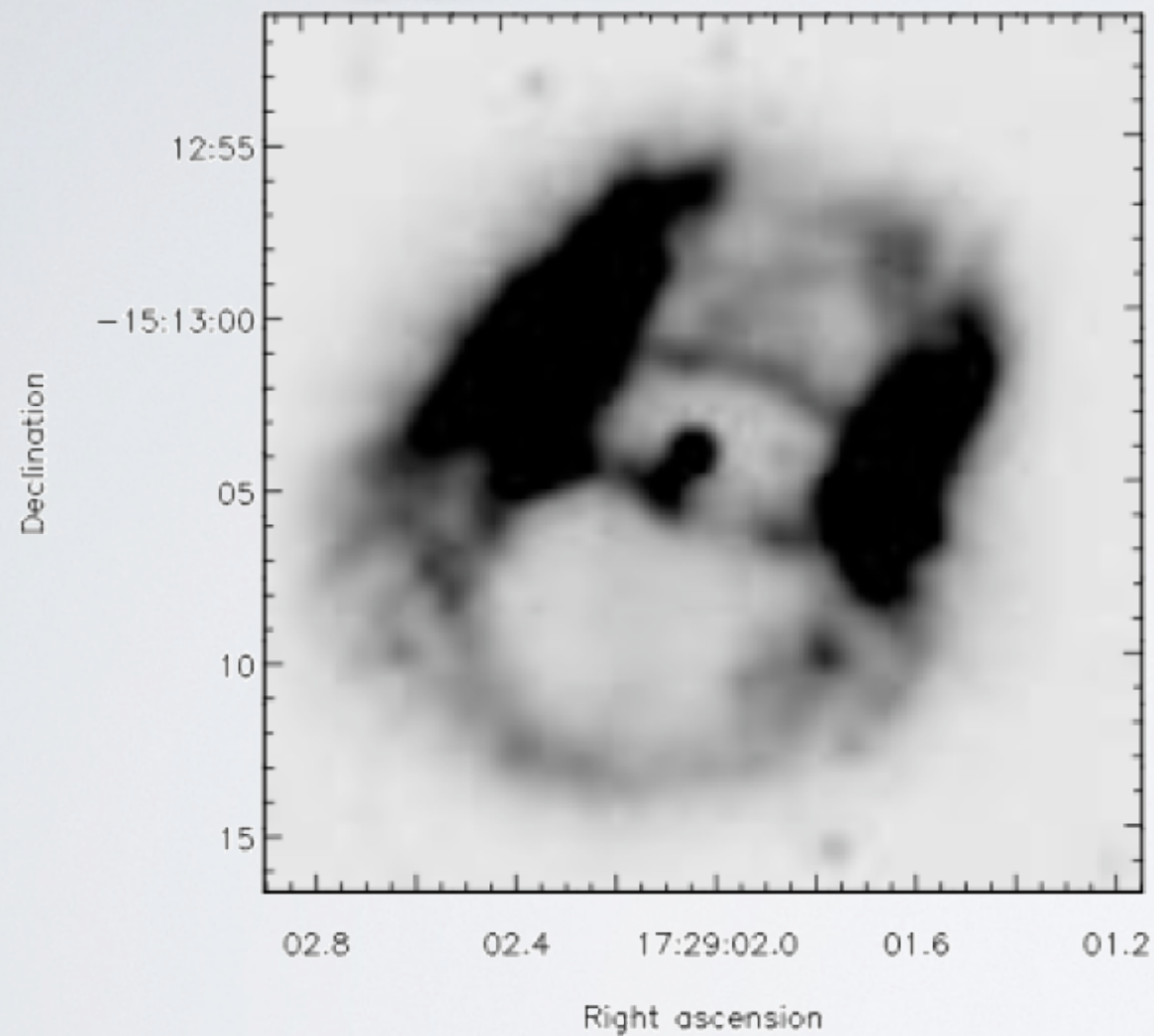


(d) 15°

3. KINEMATICS - GETTING THE MORPHOLOGY RIGHT

Jones et al. 2010, MNRAS 408, 2312

Abell 41

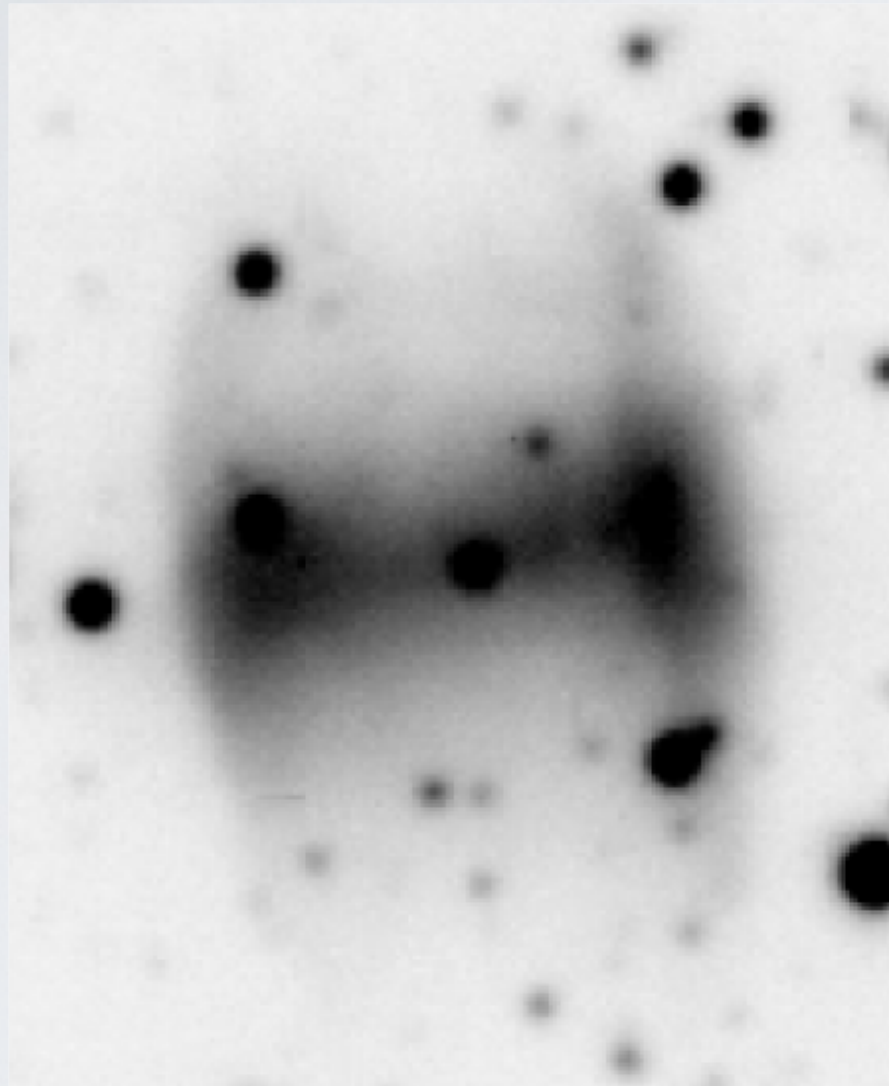


~~Elliptical~~ → Bipolar

3. KINEMATICS - GETTING THE MORPHOLOGY RIGHT

Tyndall et al. 2012, MNRAS, 422, 1804

HaTr 4

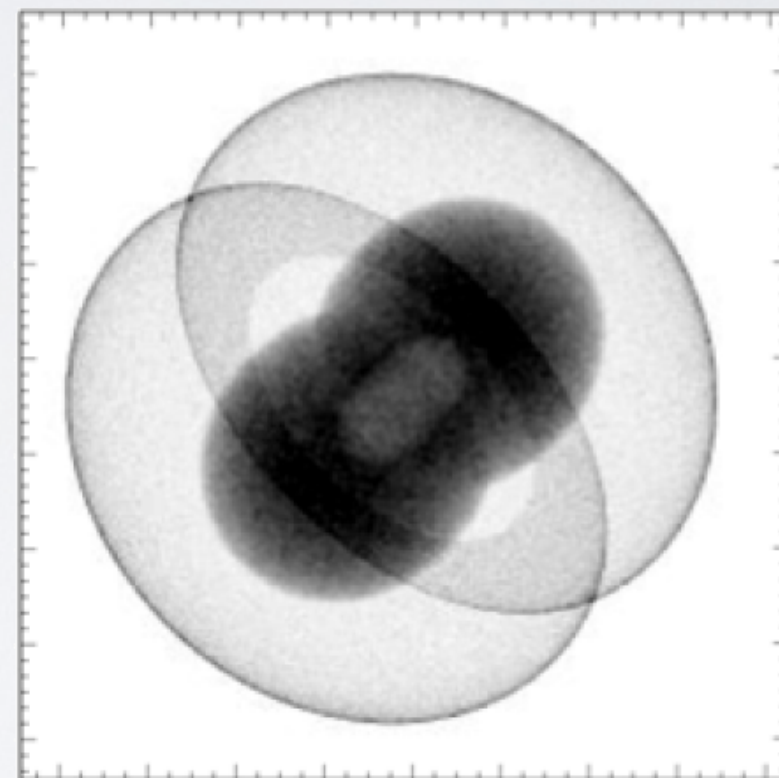


Bipolar → Elliptical

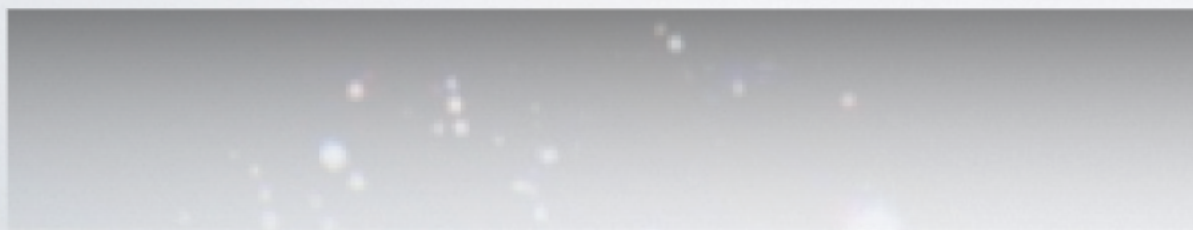
3. KINEMATICS - GETTING THE MORPHOLOGY RIGHT

Huckvale et al. 2013, MNRAS, 434, 1505

Abell 65



Multiple shell

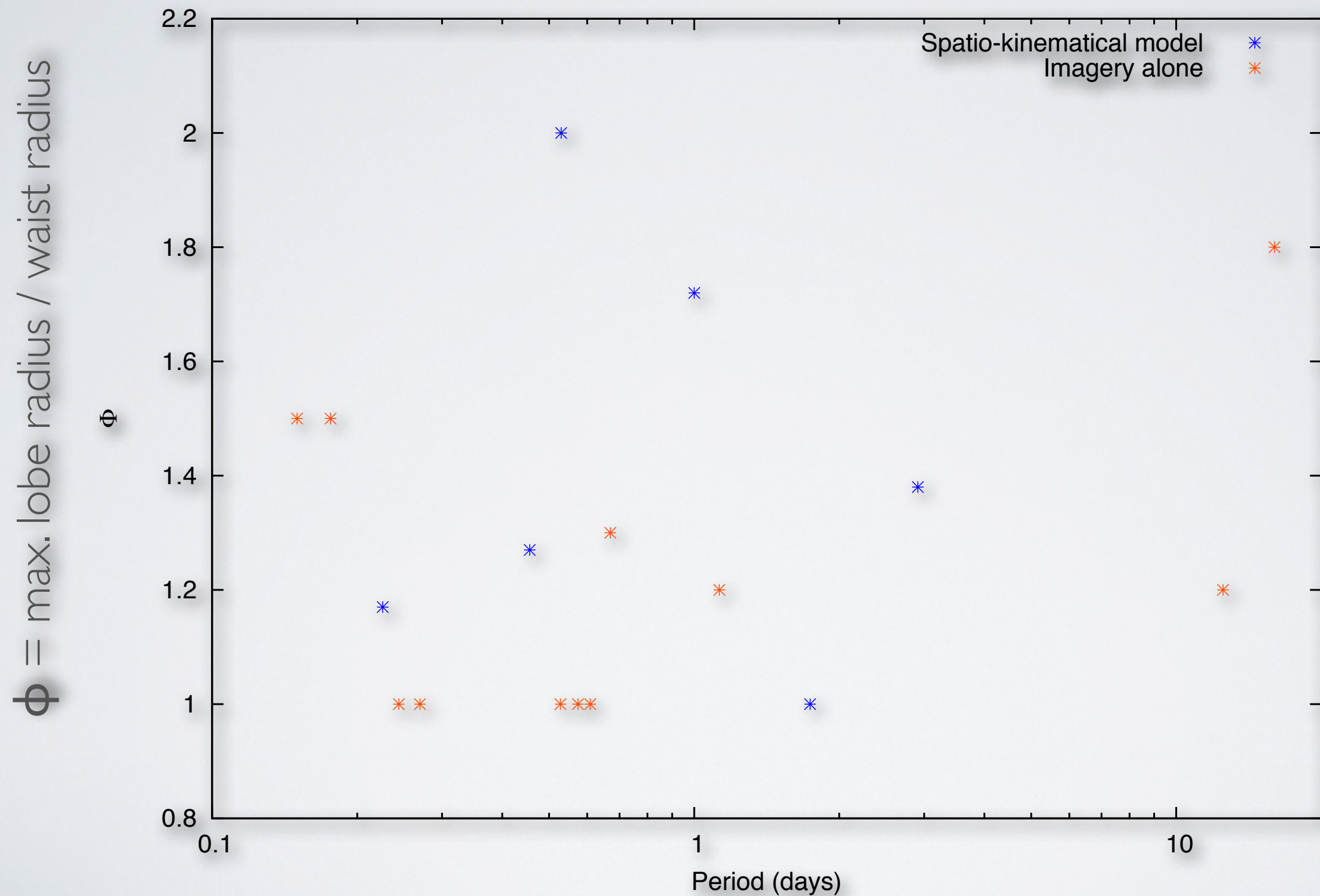


4. AND NOW WHAT? - MORPHOLOGIES

- Bipolars?
- Ellipticals?
- Irregulars?
- LIS?
- Jets?
- Rings?

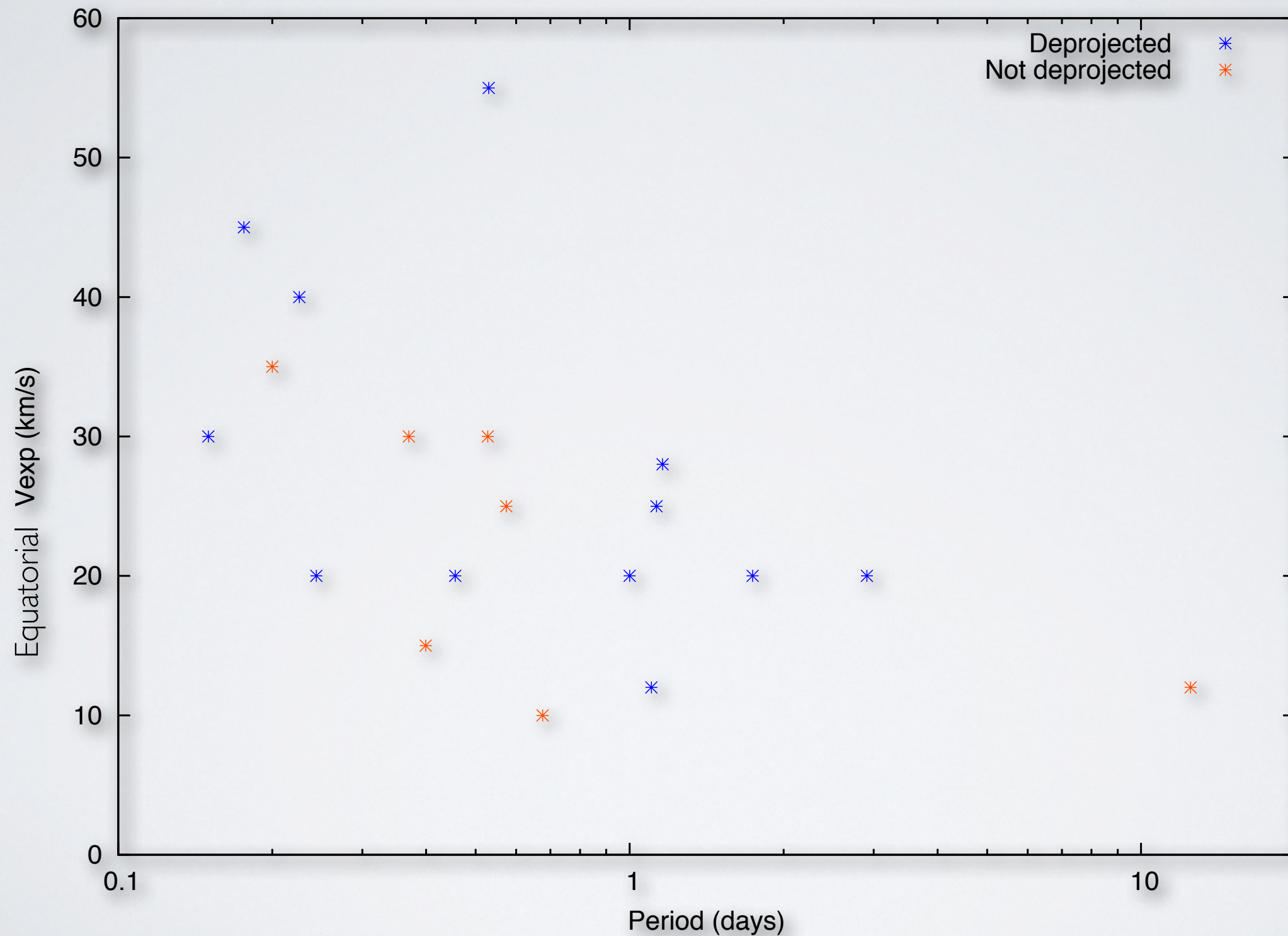
4. AND NOW WHAT? - MORPHOLOGY LINKED TO PERIOD?

Jones et al., in preparation



4. AND NOW WHAT? - EXPANSION VELOCITY LINKED TO PERIOD?

Jones et al., in preparation



5. SUMMARY

- Companion stars are a definite shaping agent for non-spherical PNe (100% rate inclination match!). Perhaps the only one?
- Morphologically disparate, but some features dominate:
 - ✓ Bipolars
 - ✓ Rings
 - ✓ Jets
 - ✓ Filaments
- No strong link to current knowledge of central binary