TABLE 10 (CONTINUED)

Code	Species	Mult.	λ(Å)	log gf	Ref.	W_{eq}	log(N/N_T)	Code	Species	Mult.	λ(Å)	log gf	Ref.	W_{eq}	log(N/N_T)
26.00	Fe I	20	3872.50	−0.93	N4	3	−4.54	26.01	Fe II	J	5018.44	−1.22	MF	77	−4.68
26.00	Fe I	41	4383.54	0.20	N4	6	−5.10	26.01	Fe II	J	5022.79	−0.02	KX	8	−4.78
26.01	Fe II	127	3845.18	−2.29	KX	8	−5.24	26.01	Fe II	J	5032.71	0.11	KX	14	−4.52
26.01	Fe II	38	3935.96	−1.86	N4	12	−4.96	26.01	Fe II	J	5035.71	0.61	KX	21	−4.79
26.01	Fe II	190	3938.97	−1.85	N4	11	−4.88	26.01	Fe II	J	5061.72	0.22	KX	7	−5.06
26.01	Fe II	28	4122.67	−3.38	N4	9	−4.99	26.01	Fe II	J	5070.90	0.24	KX	6	−5.16
26.01	Fe II	27	4273.33	−3.34	N4	8	−5.02	26.01	Fe II	J	5075.76	0.28	KX	8	−4.95
26.01	Fe II	28	4296.57	−3.01	N4	15	−5.03	26.01	Fe II	J	5082.23	−0.10	KX	5	−4.80
26.01	Fe II	27	4303.18	−2.61	N4	46	−4.47	26.01	Fe II	J	5093.58	0.11	KX	11	−4.64
26.01	Fe II	27	4351.77	−2.08	N4	35	−5.35	26.01	Fe II	J	5145.77	−0.40	KX	5	−4.49
26.01	Fe II	27	4385.39	−2.57	N4	28	−5.03	26.01	Fe II	J	5149.46	0.40	KX	14	−4.76
26.01	Fe II	27	4416.83	−2.60	N4	26	−5.05	26.01	Fe II	J	5150.49	−0.12	KX	9	−4.50
26.01	Fe II	37	4489.18	−2.97	N4	18	−4.89	26.01	Fe II	J	5169.03	−0.87	MF	81	−4.89
26.01	Fe II	37	4491.40	−2.70	N4	18	−5.16	26.01	Fe II	J	5199.12	0.10	KX	5	−5.08
26.01	Fe II	38	4508.29	−2.21	N4	32	−5.23	26.01	Fe II	J	5203.64	−0.05	KX	6	−4.81
26.01	Fe II	37	4515.34	−2.48	N4	26	−5.14	26.01	Fe II	J	5216.85	0.81	KX	17	−5.00
26.01	Fe II	37	4520.22	−2.60	N4	22	−5.15	26.01	Fe II	J	5234.62	−2.05	MF	37	−5.02
26.01	Fe II	38	4522.63	−2.03	N4	40	−5.17	26.01	Fe II	J	5247.95	0.55	N4	22	−4.55
26.01	Fe II	38	4541.52	−3.05	N4	13	−4.98	26.01	Fe II	J	5260.25	1.07	KX	30	−4.84
26.01	Fe II	37	4555.89	−2.29	N4	34	−5.10	26.01	Fe II	J	5272.40	−2.03	MF	14	−4.46
26.01	Fe II	38	4576.34	−3.04	N4	13	−5.00	26.01	Fe II	J	5276.00	−1.94	MF	34	−5.21
26.01	Fe II	37	4582.84	−3.10	N4	11	−5.07	26.01	Fe II	J	5291.67	0.58	KX	13	−4.94
26.01	Fe II	38	4583.84	−2.02	N4	51	−4.83	28.01	Ni II	11	3849.55	−1.88	KX	9	−6.54
26.01	Fe II	D	4596.02	−1.84	N4	8	−4.84	28.01	Ni II	11	4067.03	−1.29	KX	11	−7.03
26.01	Fe II	38	4620.52	−3.28	N4	8	−5.03	38.01	Sr II	1	4215.52	−0.17	WM	21	−8.35
26.01	Fe II	37	4629.34	−2.37	N4	31	−5.10	39.01	Y II	6	3950.35	−0.49	HL	21	−7.55
26.01	Fe II	186	4635.32	−1.65	N4	17	−4.77	39.01	Y II	6	3982.59	−0.49	HL	19	−7.62
26.01	Fe II	28	4666.76	−3.33	N4	10	−4.90	39.01	Y II	5	4422.58	−1.27	HL	7	−7.43
26.01	Fe II	43	4731.45	−3.13	N4	10	−5.05	39.01	Y II	22	4854.86	−0.38	HL	12	−7.59
26.01	Fe II	J	4908.15	−0.30	KX	8	−4.50	39.01	Y II	22	4900.12	−0.09	HL	24	−7.38
26.01	Fe II	J	4951.58	0.18	KX	12	−4.76	39.01	Y II	20	5119.11	−1.36	HL	4	−7.14
26.01	Fe II	J	4984.47	0.01	KX	6	−4.92	39.01	Y II	20	5200.41	−0.57	HL	14	−7.28
26.01	Fe II	J	4990.51	0.18	KX	8	−4.94	39.01	Y II	20	5205.72	−0.34	HL	20	−7.24
26.01	Fe II	J	4993.36	−3.65	MF	4	−4.95	40.01	Zr II	16	3998.95	−0.67	GB	5	−7.91
26.01	Fe II	J	5001.96	0.90	KX	28	−4.85	40.01	Zr II	42	4161.21	−0.72	BG	15	−7.26
26.01	Fe II	J	5004.20	0.50	KX	14	−4.96	80.01	Hg II	–	3983.94	−1.73	DW	50	−5.39

lated with stellar rotation. If difussion is operating in the atmospheres, probably one should expect that abundance anomalies should decrease with increasing rotational velocity. In fact, CP stars are rarely found having rotational velocities larger than ∼120 km s^{-1}. Finding a trend or its absence will be important in the context of the diffusion theory, which is the best explanation today for the spectroscopic characteristics of CP stars.

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